

1 **Supplementary materials for**

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3 Numerical and observational study of S_n -to- L_g conversion due
4 to crustal-thickening: implications for identification of
5 continental mantle earthquakes

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24 **S1. Reconstruction of undulated geometry with AxiSEM3D**

25 It is easy to use SalvusMeshLite (https://gitlab.com/swp_ethz/public/SalvusMeshLite) to generate
26 a mesh to use with AxiSEM3D. However, these meshes can only be regular, meaning that they do
27 not contain desired geometrical undulations on element boundaries (e.g. for a Cartesian domain,
28 all element boundaries will be flat and for a spherical domain all element boundaries will be
29 concentrically spherical). It also can be easy to generate an undulation file, called a geometric
30 model in AxiSEM3D. For our case, the geometric model is the Moho ramp, and we only needed
31 to specify it (distance to ramp start d , width of the ramp w , height of the ramp h , and method of
32 undulation, a sigmoid function) as a function of simulation-domain latitude, assuming the source
33 is at 90° . However, the geometrical model is quite disconnected with the finite element mesh. For
34 example, even though the mesh is spherical, the geometrical model is represented linearly in
35 simulation-domain latitude and the sampling of this geometrical model in simulation-domain
36 latitude is usually different from the element nodes in the mesh (i.e. the Gauss-Lobatto-Legendre,
37 GLL, points). Our geometrical model can be specified using many fewer points than there are
38 element nodes as our mesh is necessarily extremely dense in order to sample waves up to 5Hz.

39
40 When AxiSEM3D is given the mesh and geometrical model files, the program requires the user to
41 specify on which element boundary to deform the mesh; the surrounding area must also be
42 specified to prevent over-deforming individual elements which may cause numerical instabilities.
43 This is done for the GLL points within the deformation region by specifying a local perturbation
44 value (the amount the GLL points need to be moved in the radial direction) as a Fourier series as
45 a function of azimuthal angle to enable the representation of fully 3D models (Al-Attar & Crawford,
46 2016). The program only works with Fourier series coefficients that represent the local
47 perturbations, but never actually calculates the perturbations. This approach is also used to
48 represent all other properties in the model, even when the simulation domain is axisymmetric,
49 because at least five Fourier coefficients are needed to represent the 3D point source.

50
51 Because a user only has access to the mesh and geometrical model files, but not the intermediate
52 results from the internal computations, it is hard for a user to analyze the computed wavefield on
53 a correctly deformed mesh and to check whether the internal program has represented the

54 geometrical model correctly with the GLL points, since the sampling of the two can be quite
55 different.

56
57 To add these important capabilities, we save and output intermediate results from the program,
58 and calculate the mesh deformation externally with Fourier interpolation. The intermediate results
59 we output are the element tags for elements within the undulation region (from *Undulation.cpp*),
60 the Fourier coefficients associated with these elements for undulations (from *Undulation.cpp*) and
61 the coordinates of all GLL points (from *SE_MODEL.cpp* and *Quad.cpp*). We also output the
62 element tag for the first element whose bottom edge is the start of the undulation surface (from
63 *SE_MODEL.cpp*), and this allows us to extract the coordinates on this undulated surface exactly
64 as AxiSEM3D sees it internally. This then allows plotting such a surface as shown in [Figs. 1&6a-](#)
65 [d](#) and allows checking against the original geometrical model for accuracy of representation by
66 AxiSEM3D ([Fig. S1-1](#)).

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68 Compare computational vs. input models

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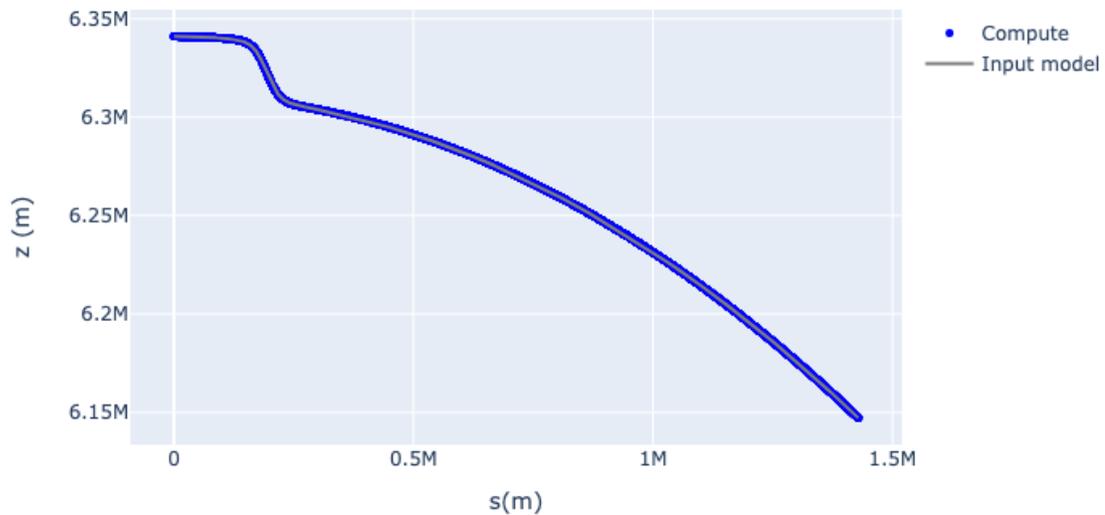
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Fig. S1-1. Checking representation of undulation by AxiSEM3D. The input geometrical model is shown as a grey line, and its representation by GLL points used for computation is shown as blue dots. The match is exact. Vertical and horizontal axes are the z , vertical, and s , radial, cylindrical coordinates measured respectively from Earth's center and from the source. M represents 10^6 . This example is shown for a Moho ramp with $d = 100$, $w = 200$ and $h = 30$ km.

81 **S2. Mesh coarsening**

82 We require the mesher to use a minimum of two elements (~10 GLL points) per wavelength in our
83 simulation domain. By default, the mesher ensures this first with the smallest wavelength, which
84 occurs in the upper crust when the shear wave velocity is the smallest. The rest of the mesh is then
85 built honoring the position of discontinuities and the background geometry and velocity models.
86 Only when the velocity at a deeper layer allows doubling the element size while still satisfying the
87 two element-per-wavelength requirement does the mesher automatically coarsen the mesh through
88 a two-refinement transition template (Fig. S2-1). For our specific case with the PREM background
89 model, this doesn't automatically happen. The consequence is that although the upper crust is
90 being sampled with 2 elements per wavelength, the deeper, faster layers are being sampled with
91 more elements per wavelength, reaching a maximum of around 2.9 below the LVZ (Fig. S2-2).
92 This number is still much lower than 4, which is needed to trigger automatic mesh coarsening.

93
94 The default mesh is not optimal in terms of computational cost, because a rather small area in the
95 computational domain (i.e. upper crust) is being sufficiently sampled, whereas the rest of the
96 domain is being over-sampled (Fig. S2-2). To combat this, we modified the shear wave velocities
97 in the background model in the upper crust and at and below the depth where we wish the element
98 size to become doubled. This depth is chosen at 70 km, i.e. 10 km below the deepest Moho in our
99 models to avoid conflict with the Moho undulation. These velocities are modified so that the
100 mesher will automatically introduce mesh coarsening at our desired depth, in such a way that when
101 the real shear velocities are replaced back, the crust, a smaller portion of the domain, will now be
102 over-sampled, whereas the majority of the simulation domain, below the coarsening depth, will be
103 sufficiently sampled with ~2 elements per wavelength (Fig. S2-2). Implementing this strategy
104 enabled us to use ~16.8% less elements for our simulations. We compare seismograms computed
105 with both coarsened and uncoarsened meshes at short and long distances to verify that our mesh-
106 coarsening technique didn't influence our results (Fig. S2-3).

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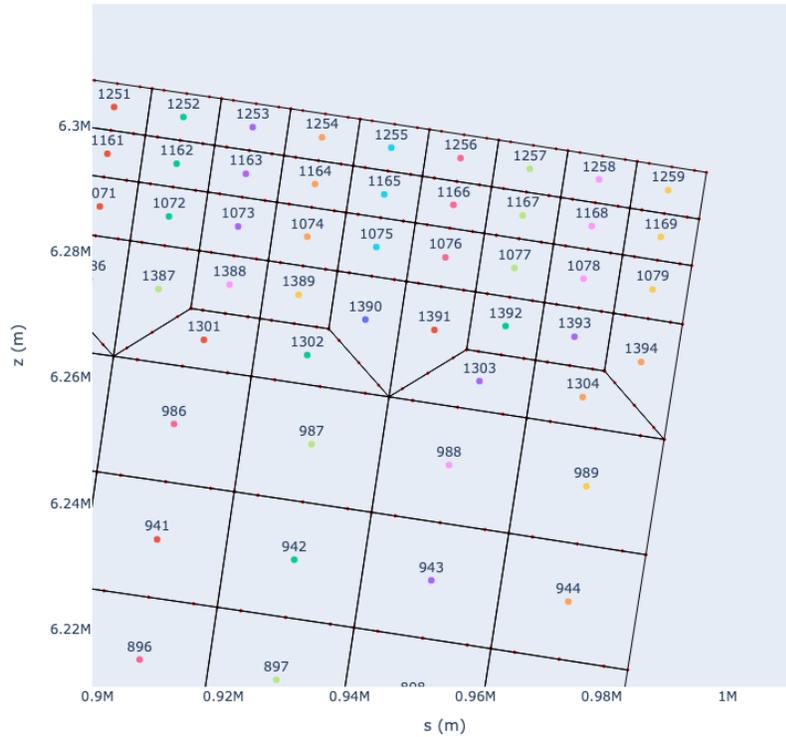


Fig. S2-1. Two-refinement transition templates. Example showing a low-resolution mesh (not used in our simulations) at an offset to demonstrate the way a mesh can be coarsened with SalvusMeshLite. The horizontal and vertical axes are the same as in Fig. S1-1. Numbers on each element are its tag number. A 2-refinement transition template consists of three elements such as elements 1387, 1388 and 1301. Faint dots on element edges are GLL points on element edges (those inside an element are not shown).

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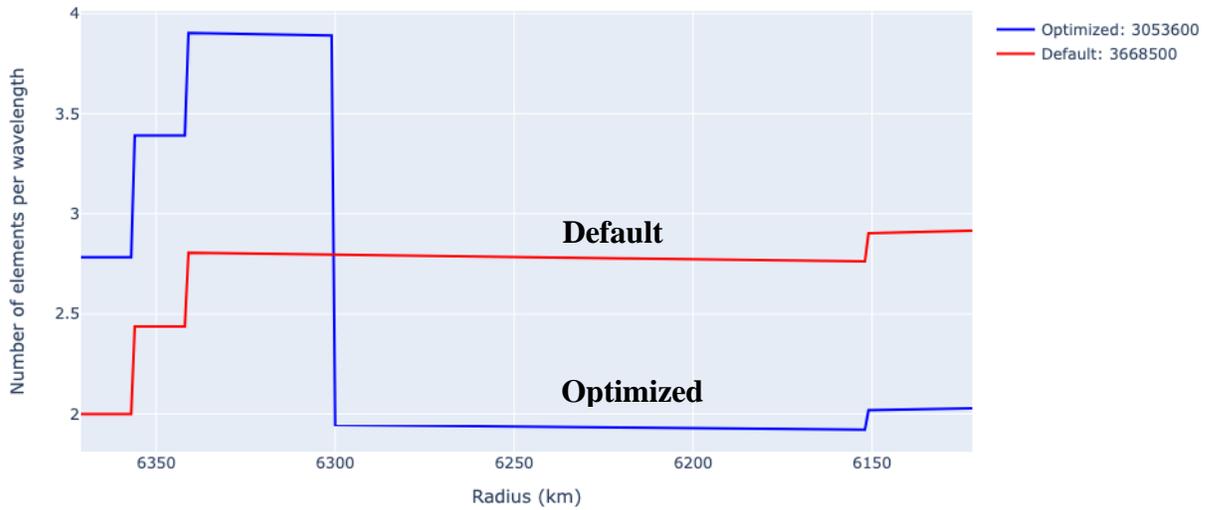


Fig. S2-2. Number of elements per wavelength for the coarsened and uncoarsened meshes. Two elements per wavelength is maintained by the default, uncoarsened mesh (red line), near the surface of the earth, with radius close to 6371 km, whereas this is honored by the optimized, coarsened mesh (blue line), at 6301 km radius, i.e. 70 km below the surface. The latter is achieved at an expense of increasing the number of elements per wavelength above this depth. The total number of elements used in the optimized and default meshes are indicated in the legend, with the optimized mesh achieving a total element number reduction of ~16.8%.

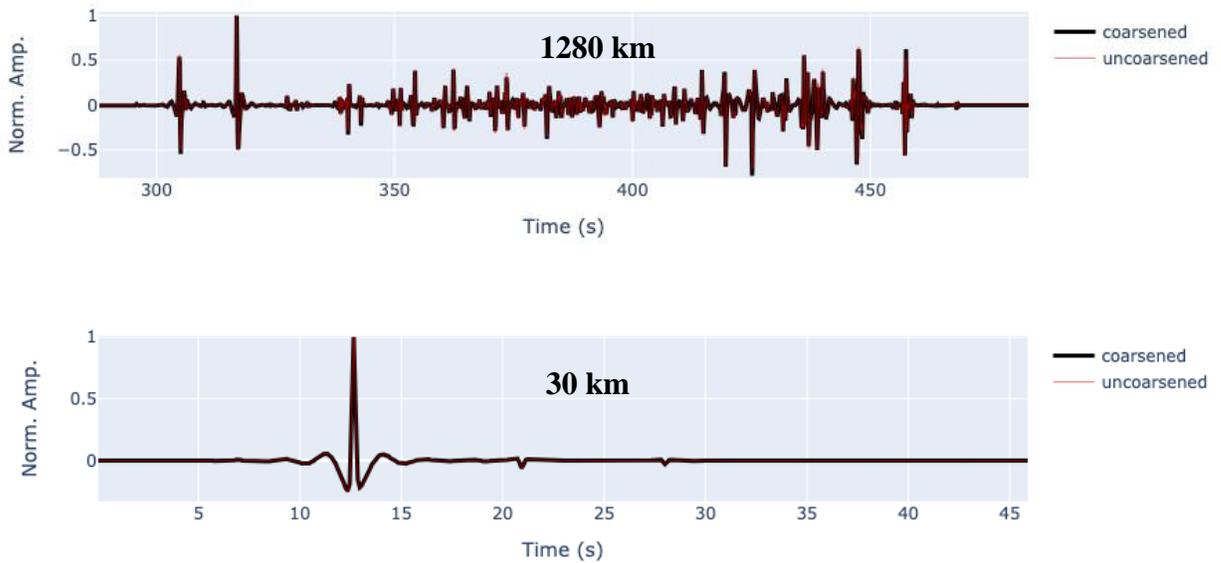


Fig. S2-3. Representative seismograms computed with the coarsened and uncoarsened meshes. Top: at epicentral distance 1,280 km. Bottom: at epicentral distance 30 km. The seismograms shown are the transverse component and are computed for the model with a Moho ramp with $d = 100$ and $w = 200$ km. Horizontal axis represents time in seconds and vertical axis is normalized amplitudes from the coarsened model (thick black line) and uncoarsened mesh (thin red line).

174 **S3. S_n and L_g windows**

175 We follow essentially the same windowing strategy as in [Wang and Klemperer \(2021\)](#), in that we
176 use the group-velocity ranges reported from regional observations and use a simple, layer-over-
177 halfspace model to calculate a zero-offset intercept for S_n , following [Barron and Priestley \(2008\)](#).
178 However, [Wang and Klemperer \(2021\)](#) maximized the S_n window by picking its end time using
179 the latest possible arrival time from a 10-km earthquake even though, for such a shallow crustal
180 earthquake, extending the S_n window this late will more likely include the earliest L_g energy than
181 any meaningful additional S_n . Instead, here we simply use the earliest and latest S_n arrival time
182 from an earthquake at the Moho as the S_n window. We suggest picking S_n and L_g windows should
183 be approached from an empirical rather than an analytical perspective, as discussed more in the
184 main text. In contrast to [Song and Klemperer \(2024\)](#) who used published earthquake depths to
185 define L_g and S_n windows as a function of hypocentral depth, we process all our data with identical
186 windows to avoid any pre-judgment and to robustly test our S_n/L_g discriminant.

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205 **S4. Additional simulation results with fixed ramp widths (w)**

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fixed width, $w=100\text{km}$

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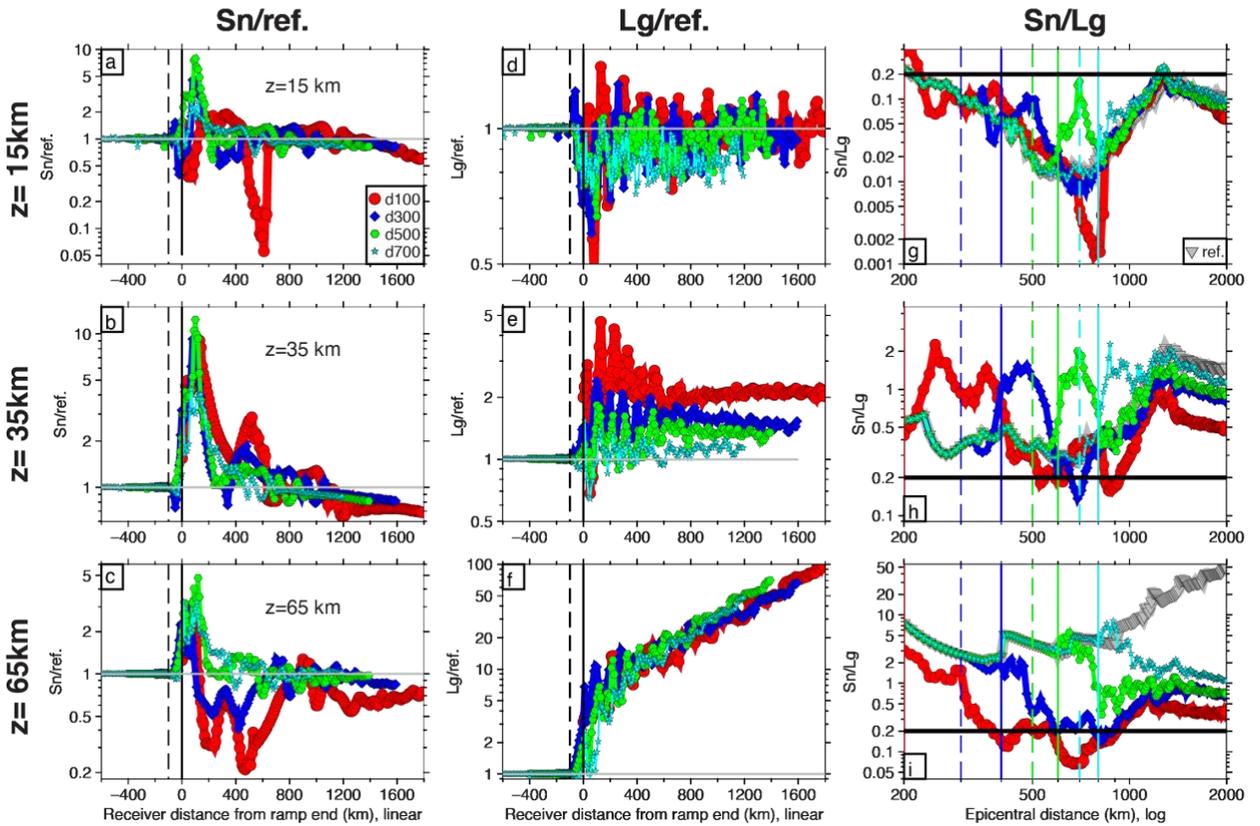
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224 **Fig. S4-1. S_n , L_g amplitudes and S_n/L_g with varying distance to ramp start d but**
 225 **fixed width $w = 100\text{ km}$.** The layout is the same as Fig. 6 except for right columns, where
 226 colored dashed lines represent the start of a ramp and colored solid lines represent the end
 227 of a ramp, for the correspondingly colored ramp models. The start of the ramp for the
 228 $d=100\text{ km}$ case is not shown and only its end is shown at 200–km.

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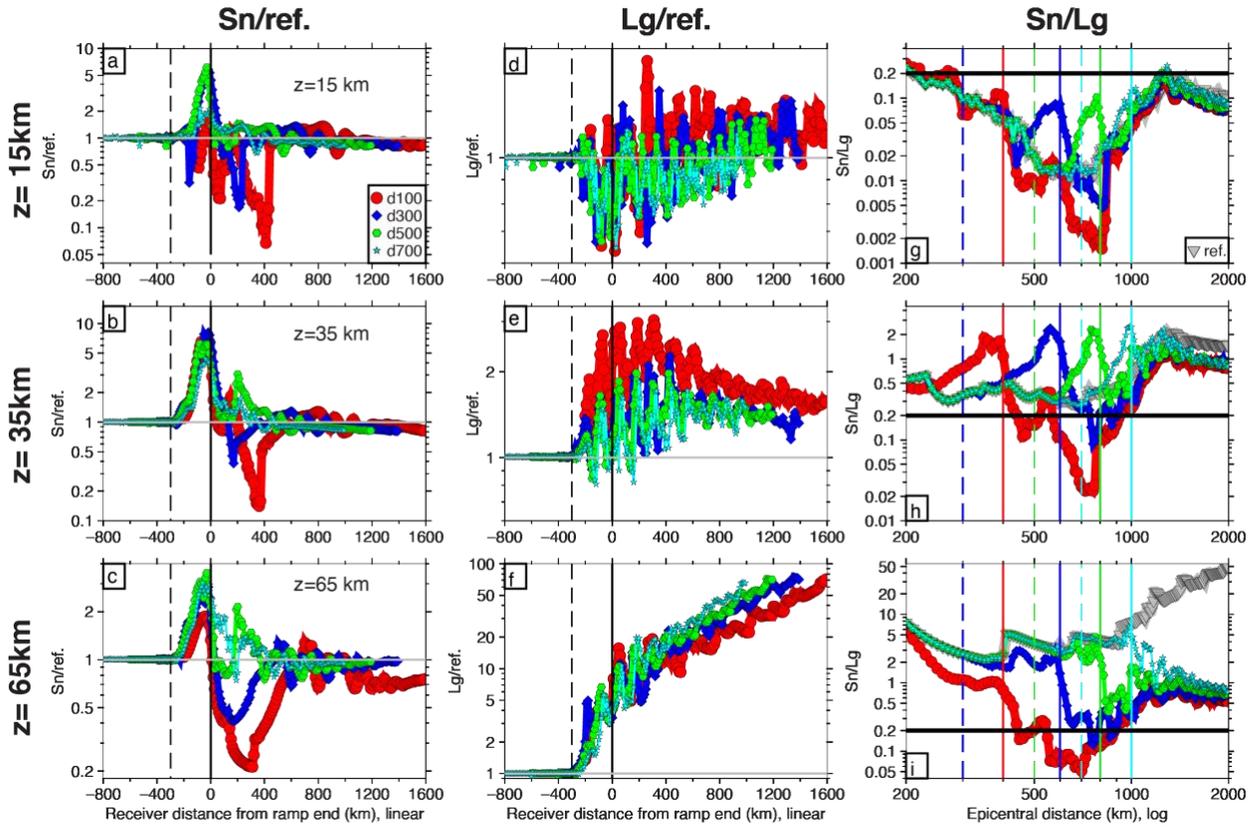
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fixed width, $w=300\text{km}$



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Fig. S4-2. S_n , L_g amplitudes and S_n/L_g with varying distance to ramp start d but fixed width $w = 300$ km. The layout is the same as Fig. S4-1.

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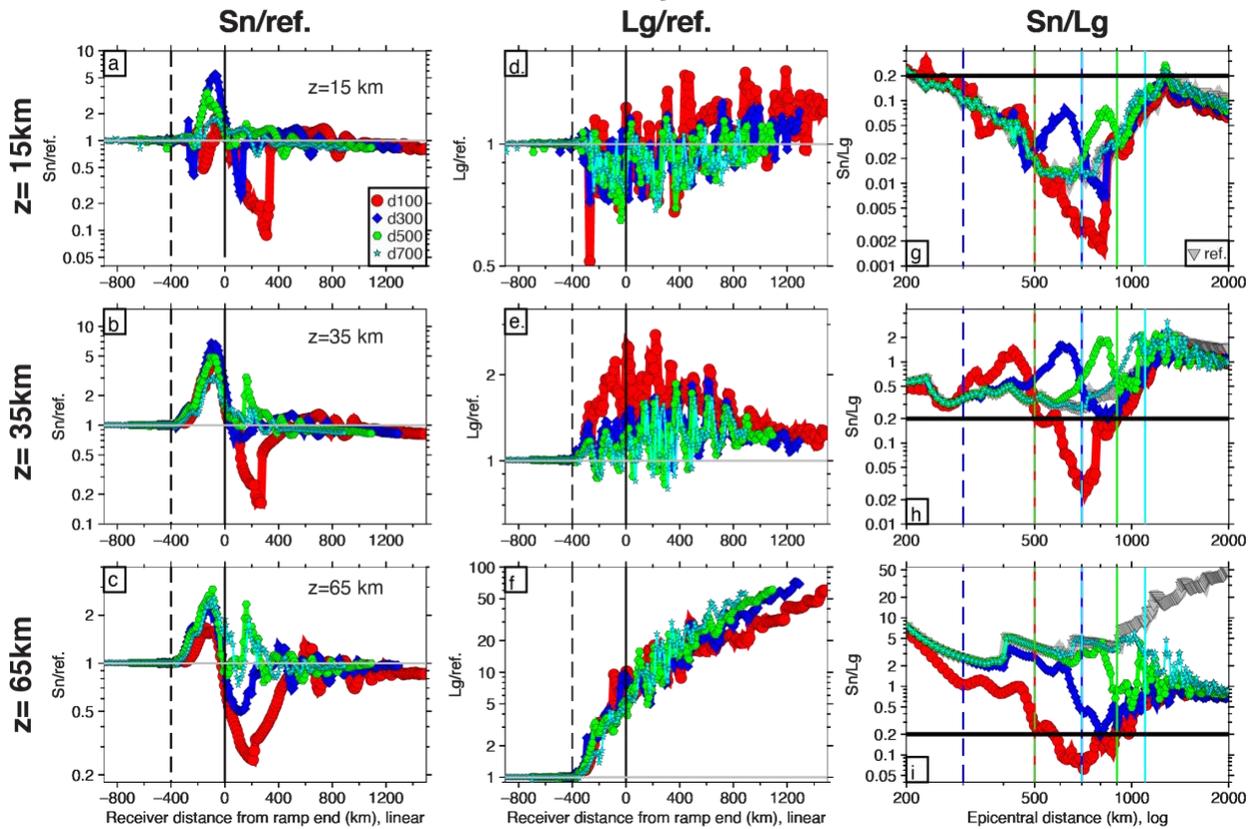
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fixed width, $w=400\text{km}$



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255 **Fig. S4-3. Sn , Lg amplitudes and Sn/Lg with varying distance to ramp start d**
 256 **but fixed width $w = 400\text{ km}$. The layout is the same as Fig. S4-1.**

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265 **S5. Additional simulation results with fixed distance to ramp (d)**

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fixed distance, $d=100\text{km}$

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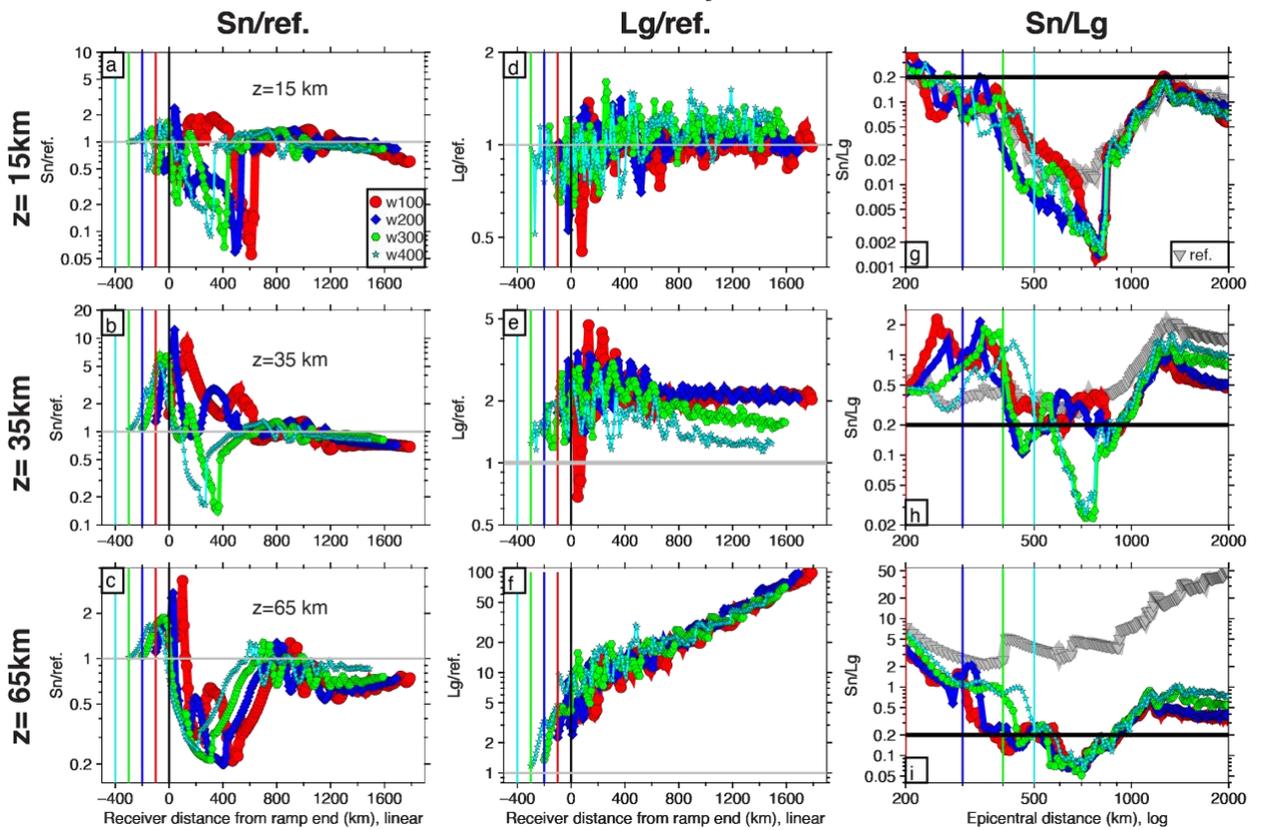


Fig. S5-1. S_n , L_g amplitudes and S_n/L_g with varying ramp width w but fixed ramp distance $d = 100$ km. The layout is the same as Fig. 9, except the start of the ramp is not shown the right column.

fixed distance, $d=500\text{km}$

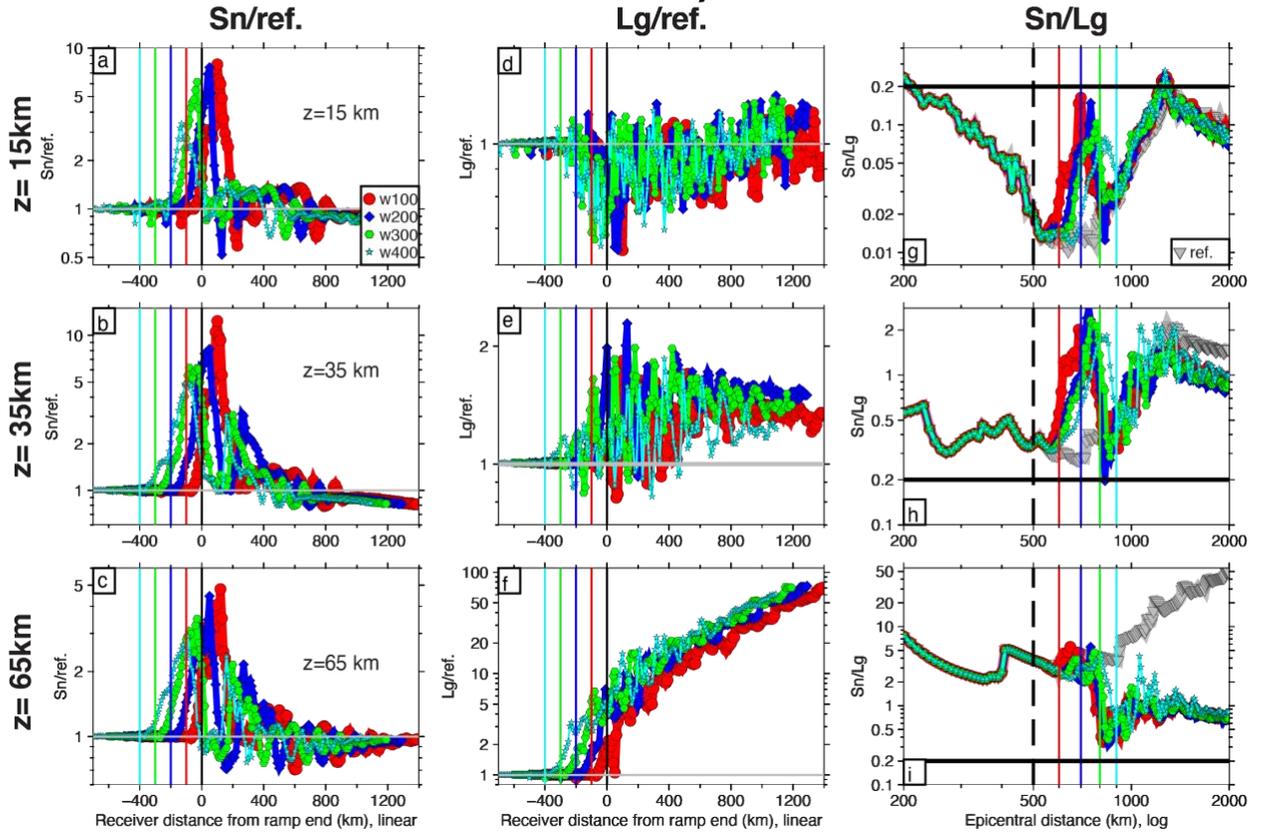


Fig. S5-2. S_n , L_g amplitudes and S_n/L_g with varying ramp width w but fixed ramp distance $d = 500$ km. The layout is the same as Fig. 9.

fixed distance, $d=700\text{km}$

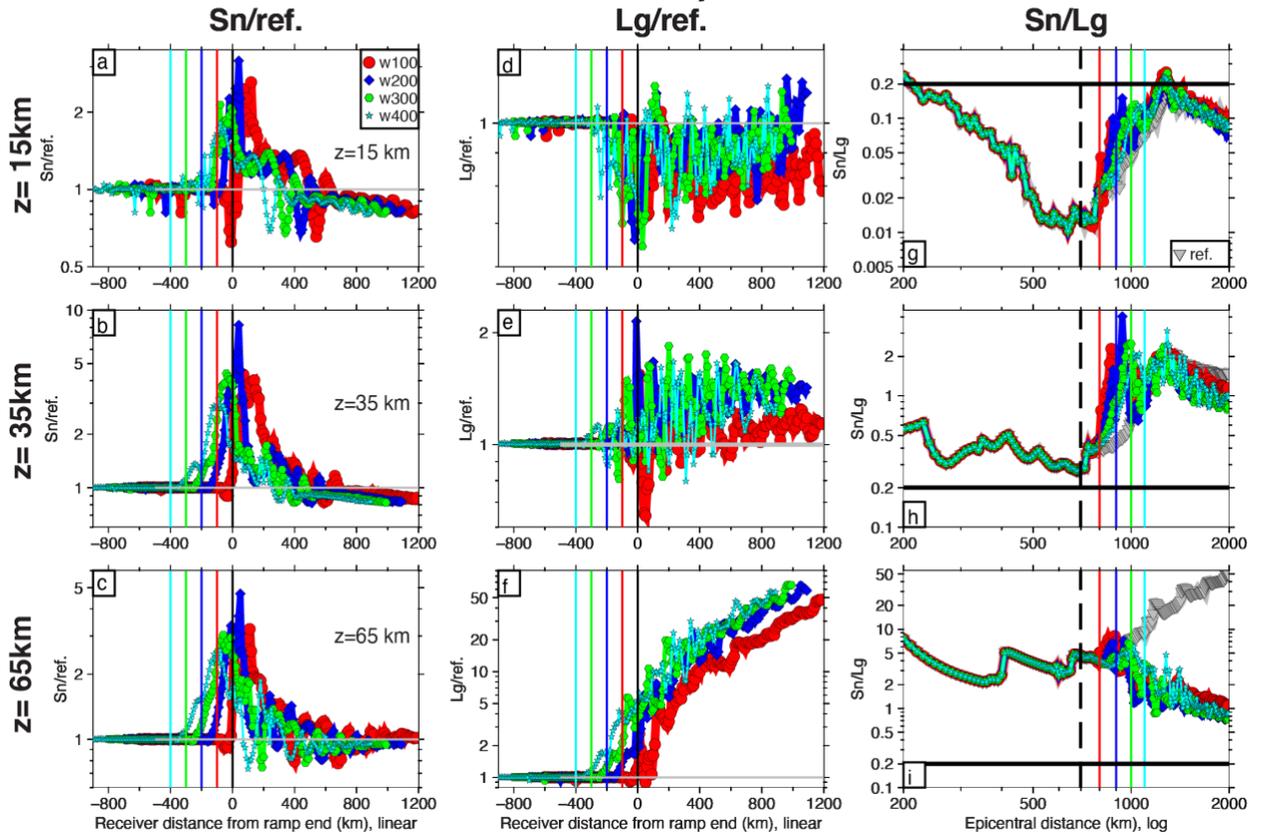


Fig. S5-3. S_n , L_g amplitudes and S_n/L_g with varying ramp width w but fixed ramp distance $d = 700\text{ km}$. The layout is the same as Fig. 9.

358 **S6. HiCLIMB events plotted against epicentral distance**

359 The southern events recorded on the HiCLIMB array are chosen so that they cover roughly the
360 same epicentral distance ranges compared to the previously studied northern events, to facilitate
361 their direct comparison. Except for a few closer stations recording event S4 and more-distant
362 stations recording S1, almost all of our recordings, both from the northern and southern events,
363 have an epicentral distance range of ~580–1100 km (Fig. S6-1). A separation of S_n/L_g , for the
364 mantle (solid symbols) and crustal (open symbols) events, can also be observed in this plotting
365 scheme (Fig. S6-1c).

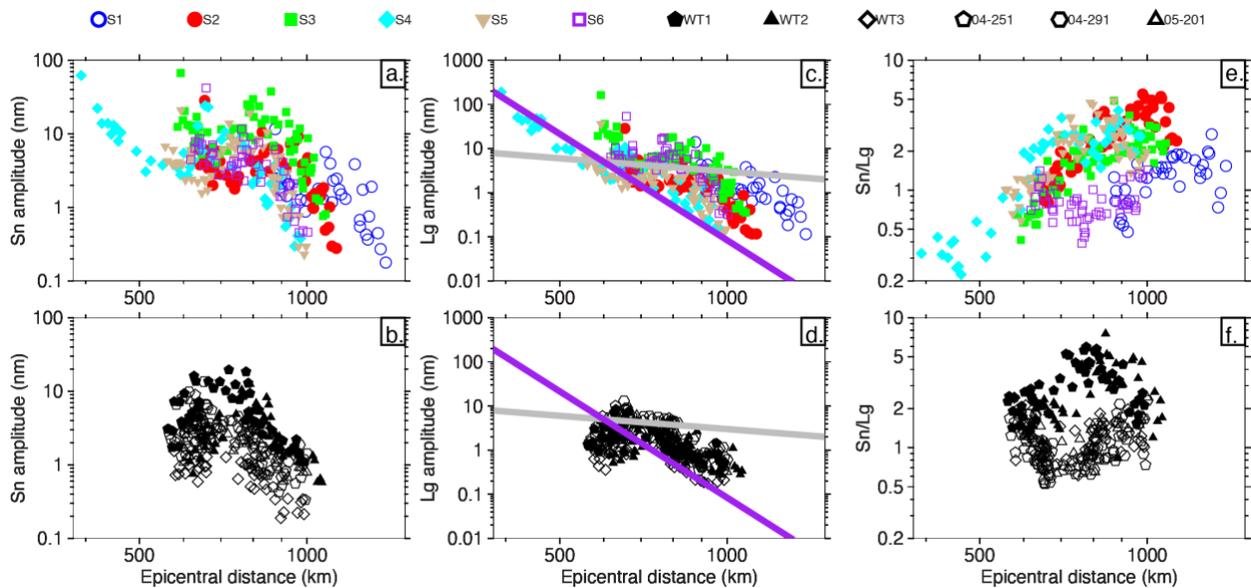
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367 This plotting scheme is ideal to investigate geometrical spreading effects in our raw data. We first
368 plot a representative L_g geometrical spreading curve with $\gamma = -1$ (Fig. S6-1c&d, grey line). It is
369 only the trend of this curve that matters but not its absolute values, i.e. it can be moved up or down
370 freely. The southern events (Fig. S6-1c, colored symbols) have a rise of L_g amplitudes at around
371 700-800 km epicentral distances, followed by a smooth amplitude decrease, i.e. the effect of the
372 attenuation region (Fig. 2a) is not visible. However, for the northern events (Fig. S6-1d, black
373 symbols), the effect of the attenuation region is visible because L_g amplitudes rise first, then
374 followed by a smooth decrease. In any case, the smooth or overall decreases of both groups of
375 events are too steep compared with the geometrical spreading model (for reference, $\gamma = -8$ is
376 plotted as a purple line in Fig. S6-1c&d, fitting the decreasing trend much better, but such a rate
377 of amplitude decrease is unreasonable for L_g geometrical spreading), and the apparent deviations
378 as discussed above are clearly correlated with station locations (Fig. 10 a-d). Hence, apparent L_g
379 geometrical spreading in real data is severely influenced by other factors such as site effects.

380
381 We then filter the S_n wavetrain at 1 Hz and at 5 Hz (with a bandpass filter from $f/\sqrt{2}$ to $\sqrt{2} f$, the
382 same way as in Yang et al. (2007), where f is the frequency of interest) and overlay the geometrical
383 spreading model at these frequencies (Fig. S6-2). All of the data start to lose amplitude at around
384 800-900 km, and the changes in amplitude are not clearly related with the interference head wave
385 phenomenon but are clearly correlated with station location and location of the attenuation zone
386 (Fig.10 a-d). Hence, we reach the same conclusion for S_n geometrical spreading as for L_g :
387 apparent spreading effects are severely influenced by local factors.

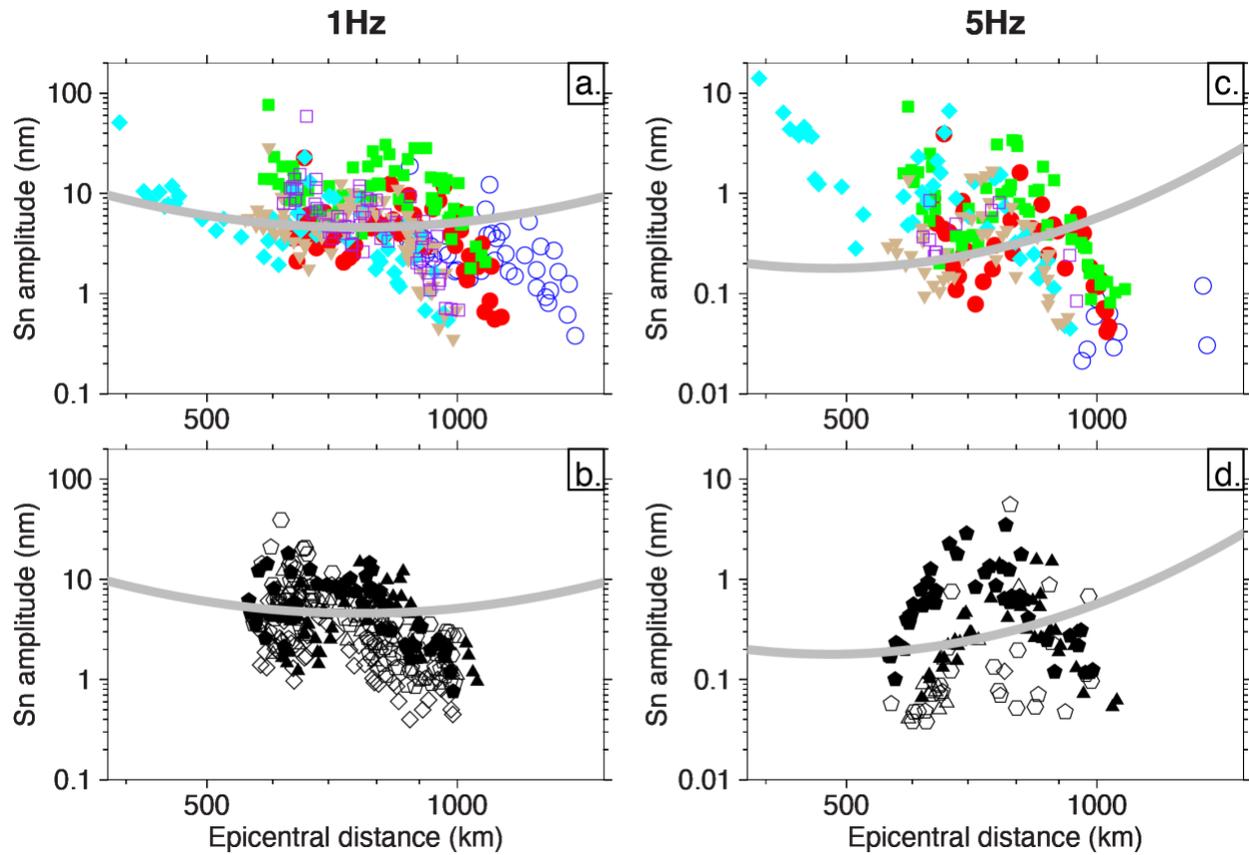
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389 These comparisons of real data and geometrical spreading models shed some lights on the apparent
 390 deviations of real data from our synthetic predictions on individual amplitudes. If the basic
 391 geometrical spreading patterns are obscured in the real data, then it is unreasonable to expect to
 392 observe the superimposed perturbations due to crossing a Moho ramp.

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 399 **Fig. S6-1. HiCLIMB data plotted against epicentral distance (log-log plots).** All symbols are the
 400 same as in the main paper for HiCLIMB events. Amplitudes are absolute without any normalization.
 401 Data are filtered from 1–5 Hz. (a)&(b) *Sn* amplitudes for northern and southern events, respectively.
 402 (c)&(d) *Lg* amplitudes for northern and southern events, respectively. Geometrical spreading model with
 $\gamma = -1$ (grey line) and $\gamma = -8$ (purple line) are shown. (e)&(f) *Sn/Lg* for northern and southern events,
 respectively. Note clear separation of solid (mantle earthquake) and open (crustal earthquake) symbols
 from ~600–1000 km. Horizontal axes for the top and bottom rows are identical and aligned, for easy
 comparison of distance ranges for our events.



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 405 **Fig. S6-2. *Sn* filtered to two frequencies.** All symbols are the same as in the main paper. Amplitudes
 406 are absolute without any normalization. (a)&(b) *Sn* at 1 Hz for northern and southern events,
 407 respectively. (c)&(d). *Sn* at 5 Hz for northern and southern events, respectively. Corresponding
 408 geometrical spreading models from [Yang et al. \(2007\)](#) are shown as grey lines.

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419 **S7. Data from the Gangdese-92 array**

Code name	Date, time	Location (N°, E°)	Magnitude (m_b)	Catalog depth (km)	Distance to ramp start (km)
G1	2012-01-01, 02:35:21	23.472, 91.834 <i>23.503, 91.804</i>	4.6	27.8±10.4 <i>15</i>	450
G2	2012-09-06, 18:27:11	25.455, 91.208 <i>25.600, 91.049</i>	4.5	45.1±8.5 <i>15</i>	245
G3	2011-09-18, 19:20:54	25.759, 91.178 <i>25.576, 91.113</i>	4.0	37.1±9.5 <i>14.2</i>	214
G4	2012-05-11, 12:41:35	26.175, 92.889 <i>26.246, 92.882</i>	5.4	43.4±? <i>43.0</i>	193

429 **Table S7-1. Events G1–G4 recorded on the Gangdese-92 array.** Event details are from
 430 USGS except Distance to Ramp, which is measured according to the description in the main
 431 paper. These events are essentially aligned with the array, so a single distance to ramp start is
 432 reported, instead of a range. The ramp width they cross is ~constant at ~160 km. Italicized
 433 hypocentral locations and depths are from the Seismological Bulletins of the Indian National
 434 Center for Seismology.

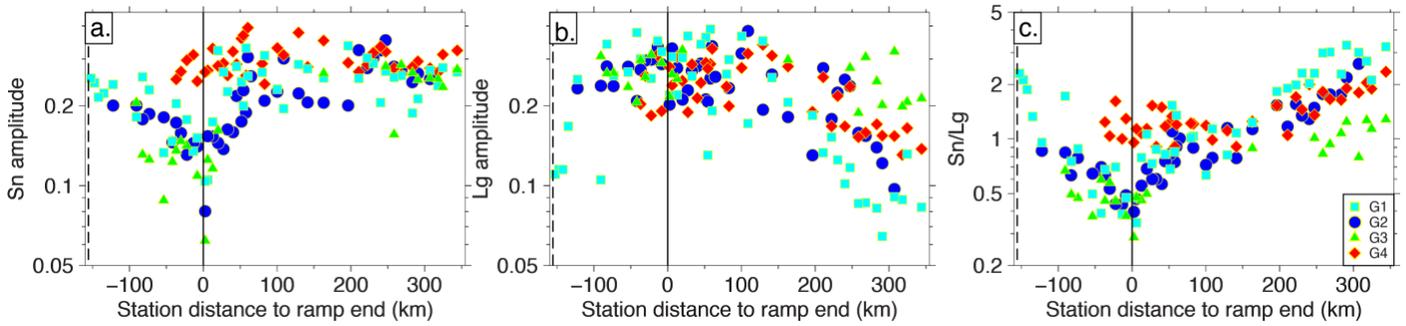
434 The S_n and L_g amplitudes and S_n/L_g amplitude ratios for events G1–G4 are shown in Fig. S7-1.
 435 Apart from S_n from G4, all other S_n and L_g amplitudes vary coherently with station locations (Fig.
 436 S7-1a&b), demonstrating likely site effects. L_g for these four events has a plateau until ~100 km
 437 beyond the end of the ramp, then at greater distances L_g amplitudes start to fall off faster than r^{-1}
 438 where r is epicentral distance. This could be a representation of enhanced L_g near the ramp end.
 439 S_n/L_g for these four events largely overlaps (as does their L_g HF/LF, Fig. 12c), and is always <2,
 440 a working S_n/L_g threshold established at the nearby PASSCAL array (Wang & Klemperer, 2021)
 441 for mantle earthquakes. Without a clear S_n/L_g baseline for collocated crustal earthquakes, we are
 442 unable to rigorously establish the depths of these events with respect to Moho, but they are very
 443 likely deep-crustal events given their catalog depths. Our inability to perform a full analysis on the
 444 Gangdese array is the reason we show its results only in supplementary materials (apart from its
 445 L_g HF/LF).

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 447 Record sections for these four events are shown in Fig. S7-2, indicating good signal-to-noise ratio.

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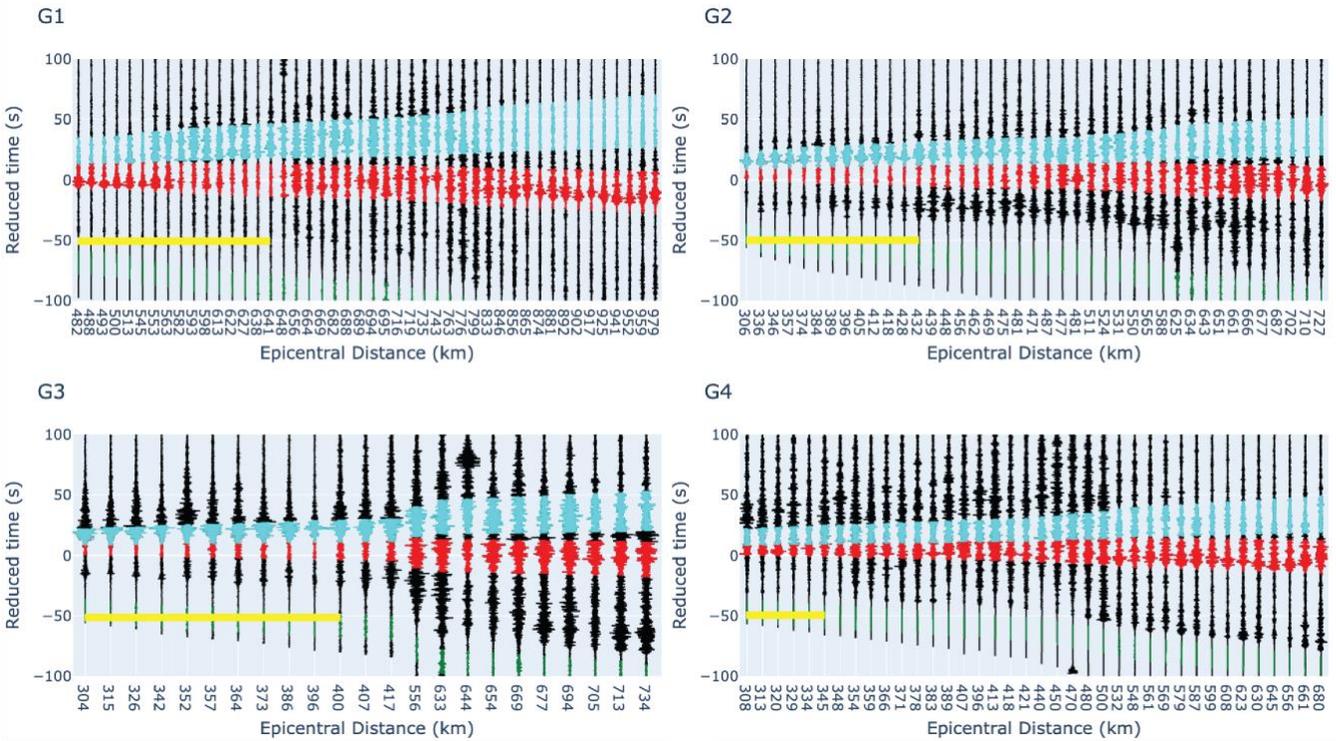
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452 **Fig. S7-1.** *Sn*, *Lg* amplitudes and *Sn/Lg* as a function of distance to ramp end for Gangdese-92
 453 events. All annotations are as in Fig. 10. This array did not extend into the *Sn* attenuation region.

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456 **Fig. S7-2.** Record sections for Gangdese-92 events. Record sections are shown with reduction
 457 velocity of 4 km/s. Green section: noise window. Red section: *Sn* window. Cyan section: *Lg*
 458 window. Horizontal yellow bar indicates stations south of YZS, i.e. within the Moho ramp region.

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461 **S8. Measuring S_n/L_g 100 km north of the end of the ramp**

462 We remeasure the percentage of $S_n/L_g > 2$ for all of the HiCLIMB events, but only use stations
463 that are at least 100 km north of YZS. This is performed to evaluate whether S_n/L_g signature is
464 stronger if we move further away from the ramp, as predicted by our synthetics and qualitatively
465 shown in [Fig. 10e](#).

466
467 The four southern events interpreted as mantle earthquakes (S2, S3, S4, S5) show a significant
468 increase in the percentage of stations that record $S_n/L_g > 2$ if we only use stations > 100 km north
469 of the end of the Moho ramp ([Fig. S8-1](#)). However, for the northern events that see no ramp, and
470 for the southern events interpreted as crustal earthquakes (S1 & S6), using only the stations > 100
471 km north of the end of the Moho ramp did not significantly change the percentage of stations
472 surpassing our threshold ([Fig. S8-1](#)). Hence we demonstrate that, using only stations > 100 km
473 distant from the ramp does increase the contrast of crustal and mantle earthquakes as measured by
474 S_n/L_g .

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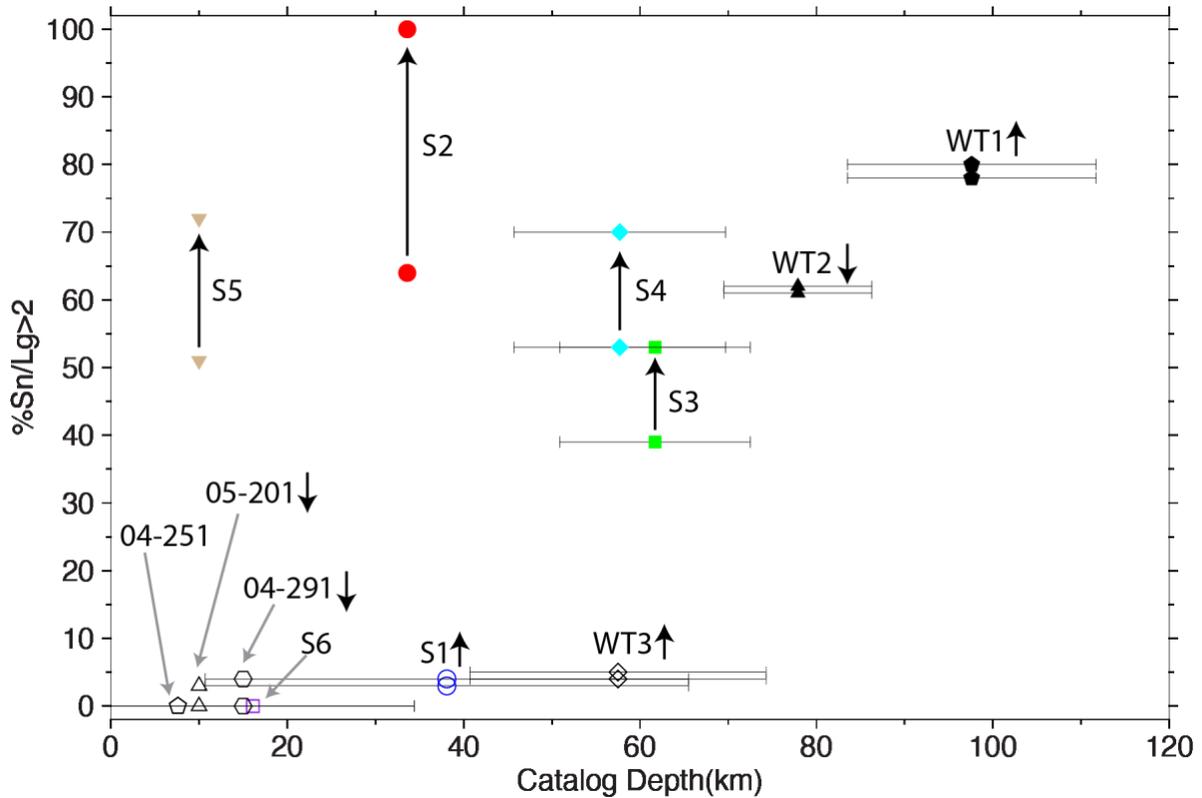
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Fig. S8-1. Percentage of stations with $S_n/L_g > 2$ for HiCLIMB events, vs. catalog depth. Symbols are the same as in Fig. 10; colored symbols S1–S6 are earthquakes in the Himalayan foreland with Moho depth likely < 45 km (Singh et al., 2015; Mitra et al., 2018), whereas black symbols are from northwest Tibet with Moho depth likely > 80 km. Horizontal axis shows catalog depths from USGS with uncertainties when available. Black arrows next to event names correspond to whether percentage of stations with $S_n/L_g > 2$ has increased or decreased from using all stations to using only stations at least 100 km north of YZS. When there is no change, no black arrow is shown. Grey arrows lead event names to event symbols.