

# Benchmark calculations of radiative forcing by greenhouse gases

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## Key Points:

- Mean clear-sky instantaneous radiative forcing by greenhouse gases is computed with six benchmark models using 100 atmospheric profiles.
- Sampling error is several times smaller than the level of disagreement among models, which is itself smaller than parameterization error.
- The impacts of clouds and stratospheric adjustment are roughly estimated; adjustments are large only for carbon dioxide and ozone.

**Abstract**

Changes in the concentration of greenhouse gases within the atmosphere lead to changes in radiative fluxes throughout the atmosphere. The value of this change, called the instantaneous radiative forcing, varies across climate models, due partly to differences in the distribution of clouds, humidity, and temperature across models, and partly due to errors introduced by approximate treatments of radiative transfer. This paper describes an experiment within the Radiative Forcing Model Intercomparison Project that uses benchmark calculations made with line-by-line models to identify parameterization error in the representation of absorption and emission by greenhouse gases. The clear-sky instantaneous forcing by greenhouse gases to which the world has been subject is computed using a set of 100 profiles, selected from a re-analysis of present-day conditions, that represent the global annual mean forcing with sampling errors of less than  $0.01 \text{ W m}^{-2}$ . Six contributing line-by-line models agree in their estimate of this forcing to within  $0.025 \text{ W m}^{-2}$  while even recently-developed parameterizations have typical errors four or more times larger, suggesting both that the samples reveal true differences among line-by-line models and that parameterization error will be readily resolved. Agreement among line-by-line models is better in the longwave than in the shortwave where differing treatments of the water vapor continuum affect estimates of forcing by carbon dioxide and methane. The impacts of clouds on instantaneous radiative forcing are estimated from climate model simulations. The adjustment due to stratospheric temperature change by assuming fixed dynamical heating. Adjustments are large only for ozone and for carbon dioxide, for which stratospheric cooling introduces modest non-linearity.

**1 Providing global-scale benchmarks for radiation parameterizations**

One of the three questions motivating the sixth phase of the Coupled Model Intercomparison Project (CMIP6, see Eyring et al., 2016) is “How does the Earth system respond to forcing?” The degree to which this question can be addressed depends partly on how well the forcing can be characterized. The measure most useful in explaining the long-term response of surface temperature is the *effective radiative forcing*, defined as change in radiative flux at the top of the atmosphere after accounting for adjustments (changes in the opacity and/or temperature of the atmosphere not associated with mean surface warming, see Sherwood et al., 2015). In support of CMIP6 the Radiative Forcing Model Intercomparison Project (RFMIP, see Pincus et al., 2016) characterizes the forcing to which models are subject using “fixed-SST” experiments (Rotstayn & Penner, 2001; Hansen, 2005) in which atmospheric composition and land use are varied but the response of sea-surface temperature and sea ice concentrations is suppressed (Forster et al., 2016).

The models participating in the previous phase of CMIP translated prescribed changes in atmospheric composition into a relatively wide range of effective radiative forcing, much of which remains even when model-specific adjustments are accounted for (e.g. Chung & Soden, 2015); initial results (Smith, Kramer, Myhre, et al., 2020) suggest that this diversity persists in CMIP6 models. Some of this variability is due a dependence on model state, especially how model-specific distributions of clouds and water vapor mask the radiative impact of changes in greenhouse gas concentrations (e.g. Huang et al., 2016). Additional variability, however, is due to model error in the *instantaneous radiative forcing*, i.e. the change in flux in the absence of adjustments, as illustrated by comparisons that use prescribed atmospheric conditions to (Ellingson et al., 1991; Collins et al., 2006; Oreopoulos et al., 2012; Pincus et al., 2015) to eliminate other causes of disagreement.

In an effort to untangle the contributions of state dependence and model error, RFMIP complements the characterization of effective radiative forcing with an assessment of errors in computations of clear-sky instantaneous radiative forcing due to greenhouse gases and aerosols. This assessment, identified as experiment *rad-irf*, is possible because there

76 is little fundamental uncertainty. Using reference “line-by-line” models, atmospheric con-  
 77 ditions and gas concentrations can be mapped to extinction with high fidelity at the very  
 78 fine spectral resolution needed to resolve each of the millions of absorption lines. Fluxes  
 79 computed with high spectral and angular resolution are then limited in precision primar-  
 80 ily by uncertainty in inputs. These reference models are known to be in very good agree-  
 81 ment with observations (e.g. Alvarado et al., 2013; Kiel et al., 2016), especially in the  
 82 absence of difficult-to-characterize clouds, given current knowledge of spectroscopy.

83 Previous assessments of radiative transfer parameterizations, focused on understand-  
 84 ing the causes of error, have examined the response to perturbations around a small num-  
 85 ber of atmospheric profiles. RFMIP builds on this long history by focusing on the global  
 86 scale relevant for climate modeling. As we explain below, we make this link by carefully  
 87 choosing a relatively small number of atmospheric states that nonetheless sample the con-  
 88 ditions needed to determine global-mean clear-sky instantaneous radiative forcing by green-  
 89 house gases. A number of reference modeling groups have provided fluxes for these sets  
 90 of conditions, providing both a benchmark against which parameterizations can be eval-  
 91 uated and information as to how reasonable choices might affect those benchmarks given  
 92 current understanding.

93 Here we describe the line-by-line calculations made for RFMIP and exploit them  
 94 to move towards benchmark estimates of the true radiative forcing to which the earth  
 95 has been subject due to increases in well-mixed greenhouse gases. We describe the con-  
 96 struction of a small set of atmospheric profiles that can be used to accurately reproduce  
 97 global-mean, annual-mean clear-sky instantaneous radiative forcing by greenhouse gases.  
 98 We summarize the reference calculations supplied to date and highlight the values of clear-  
 99 sky instantaneous radiative forcing for a range of changes in atmospheric composition  
 100 relative to pre-industrial conditions. We show that sampling error from the small set of  
 101 profiles is small enough that small differences among line-by-line calculations can be re-  
 102 solved, while variance among reference models is still less than even modern parameter-  
 103 ized treatments, suggesting the the experiments can identify true variability across line-  
 104 by-line models and parameterization error. We then cautiously extend these benchmark  
 105 estimates towards more useful estimates that include the impact of clouds and adjust-  
 106 ments.

## 107 2 Making global-mean benchmarks practical

108 Increasing computing power and more flexible software have made large-scale line-  
 109 by-line calculations increasingly practical. Indeed RFMIP effort to diagnose errors in in-  
 110 stantaneous radiative forcing by aerosols applies line-by-line modeling at relatively low  
 111 spectral resolution (Jones et al., 2017) to eight global snapshots for each participating  
 112 model. Errors in global mean, annual mean clear-sky instantaneous radiative forcing by  
 113 greenhouse gases, however, can be assessed with a much more parsimonious set of at-  
 114 mospheric conditions. This is because temporal variations of temperature and water va-  
 115 por are relatively slow and have a modest impact on the sensitivity of flux to changes  
 116 in greenhouse gas concentrations. Many previous calculations (see Etminan et al., 2016,  
 117 for a recent example), in fact, estimate global mean, annual mean values using just two  
 118 or three profiles, based on work in the 1990s showing that even such simple representa-  
 119 tions of latitudinal variability are sufficient to constrain flux changes at the tropopause  
 120 to within about a percent (Freckleton et al., 1998; Myhre et al., 1998).

121 Here we describe the construction of a set of atmospheric profiles designed to de-  
 122 termine *error* in global-mean clear-sky instantaneous radiative forcing, obtained using  
 123 a reference model on a very large number of atmospheric and surface conditions to de-  
 124 termine this forcing and choosing a subset of these conditions that minimizes the sam-  
 125 pling error across a range of measures. As we demonstrate below, the same set of pro-

**Table 1.** Perturbations around present-day (PD) conditions used to identify representative profiles. These are similar to, but not the same as, the perturbations used in RFMIP experiment *rad-irf* for reasons described in the text. Perturbations are applied to each profile drawn from ERA-Interim profile set. Carbon dioxide concentrations are relative to a pre-industrial (PI) volume mixing ratio of 278 ppmv. GHG refers to well-mixed greenhouse gases. Temperature T and relative humidity RH perturbations (12, 13) use the average of two models from the CMIP5 archive (GFDL-CM3 and GFDL-ESM2G) with relatively low and high climate sensitivities, respectively.

	Perturbation
1	PI $0.5 \times \text{CO}_2$
2	PI $2 \times \text{CO}_2$
3	PI $3 \times \text{CO}_2$
4	PI $8 \times \text{CO}_2$
5	PI $\text{CO}_2$ (278 ppmv)
6	PI $\text{CH}_4$ (0.722 ppmv)
7	PI $\text{N}_2\text{O}$ (0.273 ppmv)
8	PI HFC (all HFC at zero)
9	PI $\text{O}_3$ (from CMIP6 PI ozone file)
10	PD +4K temperature, no $\text{H}_2\text{O}$ change
11	PD +20% humidity
12	PI T, RH, $\text{O}_3$ , GHG
13	2095 RCP8.5 T, RH, $\text{O}_3$ , GHG
14	PI $\text{O}_3$ , GHG
15	PI $\text{O}_3$ , GHG, but PI $4 \times \text{CO}_2$
16	2095 Avg Sens RCP4.5 $\text{O}_3$ , GHG
17	2095 Avg Sens RCP8.5 $\text{O}_3$ , GHG

126 files also provides an accurate sample of the parameterization or approximation error in  
 127 radiative forcing.

## 128 **2.1 Computing global-mean, annual mean radiative fluxes and flux per-** 129 **turbations**

130 We characterize the range of conditions in the present-day atmosphere using a sin-  
 131 gle year (2014) of the ERA-Interim reanalysis (Dee et al., 2011). We sample tempera-  
 132 ture, pressure, specific humidity, ozone mixing ratios, and surface temperature and albedo  
 133 on a  $1.5^\circ$  grid every 10.25 days. Sampling at high latitudes is reduced to maintain roughly  
 134 equal area weighting. Concentrations of other greenhouse gases ( $\text{CO}_2$ ,  $\text{CH}_4$ ,  $\text{N}_2\text{O}$ , HCFCs  
 135 22 and 134a, CFCs 11, 12, and 113, and  $\text{CCl}_4$ ) use 2014 values from NOAA greenhouse  
 136 gas inventories and are assumed to be spatially uniform. We assume that these 823,680  
 137 profiles adequately represent global-mean, annual-mean clear-sky conditions.

138 We apply a series of 17 perturbations (detailed in Table 2.1) to these conditions,  
 139 including varying concentrations of greenhouse gases (especially  $\text{CO}_2$ ), temperature, and  
 140 humidity. Some temperature perturbations include spatial patterns obtained from cli-  
 141 mate change simulations made for CMIP5. The perturbations are intended to sample  
 142 error across a wide range of conditions. The perturbations are similar to, but not quite  
 143 the same as, those used by the final RFMIP experiments in Section 3, because the RFMIP  
 144 protocol was not fully established when we performed these calculations.

Our aim is to reproduce the mean of a set of reference fluxes, fully resolved in space and time and across the electromagnetic spectrum, computed for present-day conditions and each perturbation. The fluxes are computed using the UK Met Office SOCRATES (Suite Of Community RAdiative Transfer codes based on Edwards & Slingo, 1996) configured as a narrow-band model with a very high-resolution  $k$ -distribution with 300 bands in the longwave and 260 bands in the shortwave (Walters et al., 2019). This configuration agrees quite well with line-by-line models (e.g Pincus et al., 2015) and is one of the benchmark models described in Section 3.1. The spectral overlap of gases is treated with equivalent extinction with corrected scaling. Clouds and aerosols are not considered, consistent with the protocol for RFMIP experiment *rad-irf*.

We also compute fluxes for these sets of atmospheric conditions with an approximate model: RRTMG (Mlawer et al., 1997; Iacono et al., 2000), which is based on somewhat older spectroscopic information and so is expected to have errors with a potential dependence on atmospheric state.

## 2.2 Choosing a set of globally-representative profiles

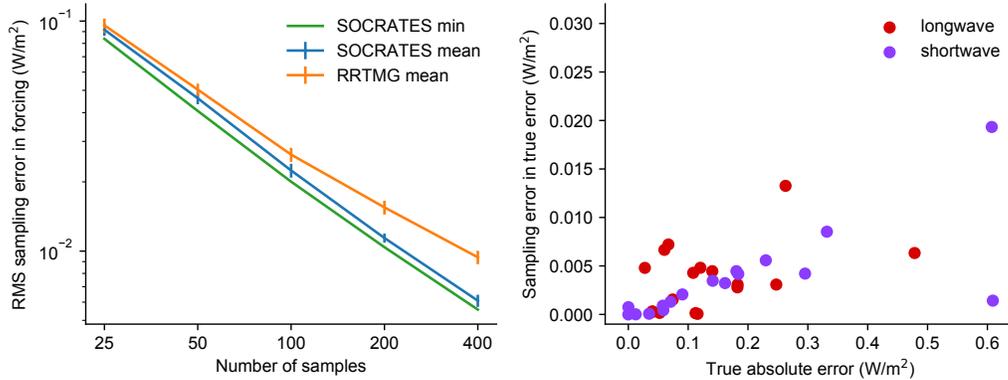
We seek a small subset of atmospheric profiles that minimizes sampling error in the global, annual mean obtained from the full calculation. To identify such a subset we must quantify what we mean by “best” by defining a cost or objective function with which to measure sampling error. Because the goal of RFMIP is to establish accuracy in calculations of radiative forcing, our objective function  $O$  is defined in terms of the change in flux between each of the 17 perturbations and present-day conditions. (For perturbations in which the only change is to greenhouse gas concentrations this quantity is precisely the instantaneous radiative forcing.) The objective function includes errors in changes of upward flux at the top of the atmosphere and downward flux at the surface as well as changes in flux divergence above and below the tropopause (the level of which is determined by Wilcox et al., 2011); each quantity is computed for both longwave and shortwave fluxes. We guard against compensating errors related to temperature, humidity, and surface albedo and emissivity by further considering 9 roughly equal-area latitude bands centered on the equator. We choose an  $l^2$  norm so that

$$O = \left[ \frac{1}{N_{\text{lat}} N_{\text{pert}} N_{\text{quant}}} \sum_l \sum_p \sum_q \left( \widehat{\Delta F}_{l,p,q} - \Delta F_{l,p,q} \right)^2 \right]^{1/2} \quad (1)$$

where  $\Delta F_{l,p,q}$  describes the average change in flux or flux divergence, as computed with the reference model over the full set of profiles, between perturbation  $p$  and present-day conditions in latitude band  $l$  for quantity  $q$ , and  $\widehat{\Delta F}_{l,p,q}$  the sampled estimate of the same quantity. The objective function includes the four flux quantities for both longwave and shortwave fluxes ( $N_{\text{quant}} = 8$ ).

We identify optimal subsets of profiles from within the complete set using simulated annealing (Kirkpatrick et al., 1983). Because the optimization is stochastic we perform 25 independent optimizations for each of a range of subset sizes. We save the realization with the lowest value of  $O$  although this choice has little impact as the standard deviation across realizations is small (roughly 6% of the mean sampling error), so that the sampling error in the best realization is only about 10% smaller than the mean (Figure 1). Simulated annealing produces sampling errors substantially lower than purely random sampling (by a factor of 19 for 100 profiles, not shown). The choice of profiles is reasonably robust to the choice of model: sampling error in the independent estimate of mean radiative forcing with RRTMG is only modestly larger (15% for 100 profiles) than for calculations with the narrow-band configuration of SOCRATES.

Profiles chosen to minimize sampling error in mean radiative forcing also provide accurate estimates of parameterization error  $\mathcal{E} = \Delta \widehat{F} - \Delta F$  in that forcing, where  $\Delta \widehat{F}$



**Figure 1.** Left: values of the cost function  $O$ , an aggregate measure of error across regions, changes in atmospheric conditions, and measures of flux (Eq. 1) as a function of the number of optimal profiles. The simulated annealing method used to chose the profiles is stochastic; the mean and standard deviation across realizations is shown along with the value of sample error from the best-fit realization used in further calculations. The choice of profiles based on reference radiative transfer calculations (“SOCRATES”) is robust, producing only modestly larger sampling errors for approximate calculations (“RRTMG”). Right: Absolute value of the sampling error  $\hat{\mathcal{E}} - \mathcal{E}$  in estimates of the approximation error  $\mathcal{E} = \Delta\tilde{F} - \Delta F$  sought by RFMIP. Errors shown are for the mean of 100 samples representing the global, annual mean, for changes in upwelling longwave flux at the top of the atmosphere (red) and downwelling shortwave flux at the surface (purple) from 17 perturbations. Parameterization errors range from 0 to about 0 to 0.6  $\text{W m}^{-2}$  in the global, annual mean; sampling error is almost always less than  $0.01 \text{ W m}^{-2}$ .

192 is a computation made with an approximate model. Fig. 1 shows the sampling error  $\hat{\mathcal{E}} -$   
 193  $\mathcal{E}$  in estimates of the global, annual mean parameterization error for RRTMG compared  
 194 to high-resolution SOCRATES calculations for the 17 perturbations used to develop the  
 195 profile samples. True absolute errors from RRTMG range from near 0 to  $0.6 \text{ W m}^{-2}$   
 196 in the global, annual mean; sampling error in these estimates is almost always less than  $0.01$   
 197  $\text{W m}^{-2}$ .

198 The RFMIP protocol uses the set of 100 profiles with the lowest value of the ob-  
 199 jective function  $O$ . As a consequence of optimizing the sampling for radiative forcing,  
 200 fluxes for any individual state including the present-day baseline are themselves subject  
 201 to sampling errors: global mean insolation in our sample, for example, is  $335.1 \text{ W m}^{-2}$   
 202 (c.f. the true mean of  $\sim 1361/4 = 340.25 \text{ W m}^{-2}$ ). In addition, using a single set of pro-  
 203 files for both longwave and shortwave calculations means that the sun is below the hori-  
 204 zon for roughly half the set of profiles.

### 205 3 Radiation calculations with reference models

206 Experiment *rad-irf* requests fluxes for these 100 profiles and for 17 perturbations  
 207 around present-day conditions, including changes in greenhouse gas concentrations, tem-  
 208 perature, and humidity (see tables 3 and 4 in Pincus et al., 2016). Below we focus on  
 209 the thirteen experiments in which gas concentrations alone are changed.

### 3.1 Contributions and variants

To date six benchmark models have contributed results: ARTS 2.3 (Buehler et al., 2018), provided by the University of Hamburg; LBLRTM v12.8 (Clough et al., 2005), provided by Atmospheric and Environmental Research; the SOCRATES narrow-band configuration described in Sec. 2.1, provided by the UK Met Office; the Reference Forward Model (Dudhia, 2017), provided by the NOAA Geophysical Fluid Dynamics Lab; GRTCODE, a new line-by-line code developed at GFDL; and 4AOP (Scott & Chédin, 1981; Chérut et al., 1995), provided by the Laboratoire de Météorologie Dynamique. Half the models use spectroscopic information from HITRAN 2012 (Rothman et al., 2013), while GRTCODE results are based on HITRAN 2016 (Gordon et al., 2017), 4AOP uses GEISA 2015 (Jacquinet-Husson et al., 2016), and LBLRTM employs the aer.v\_3.6 line file, which is based on HITRAN 2012 but includes small changes to improve comparisons with select observations. With one exception noted below the models use variants of the MT\_CKD continuum (Mlawer et al., 2012).

These six models provide eighteen sets of longwave fluxes and nine sets of shortwave fluxes. This multiplicity arises because some models provided calculations for slightly different sets of greenhouse gases, called “forcing variants” within CMIP and RFMIP, and/or slightly different model configurations (“physics variants”).

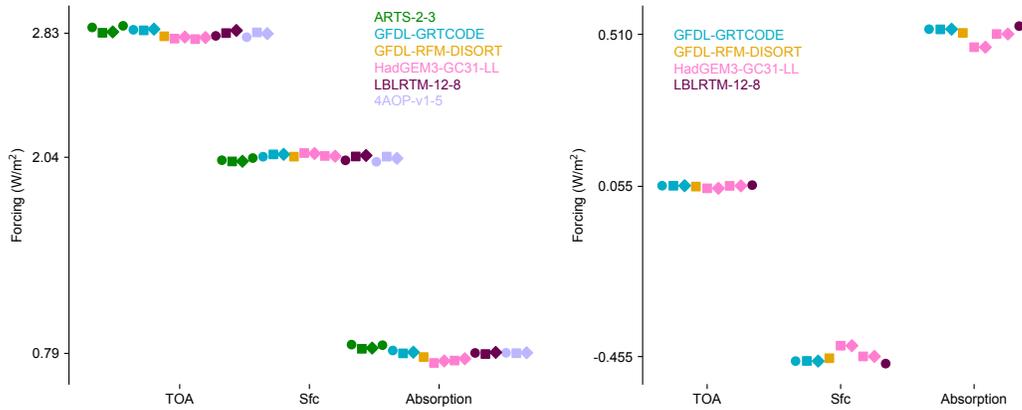
Climate models participating in CMIP6 may specify well-mixed greenhouse concentrations using one of three forcing variants described by Meinshausen et al. (2017): using some or all of the 43 greenhouse gases provided in the forcing data set; by prescribing CO<sub>2</sub>, CH<sub>4</sub>, N<sub>2</sub>O, CFC-12, and an “equivalent” concentration of CFC-11 to represent all other gases; or using CO<sub>2</sub>, CH<sub>4</sub>, N<sub>2</sub>O, and equivalent concentrations of CFC-11 and HFC-134a. (Concentrations of water vapor and ozone are drawn from reanalysis, as described in Sec. 2.1.) Some models provided results for more than one of these forcing variants.

In addition, some models provided calculations with slightly reconfigured models. ARTS 2.3 does not normally include CO<sub>2</sub> line mixing but provided a second physics variant that did so. High spectral resolution calculations with SOCRATES are themselves considered a second physics variant of the lower-resolution calculations made during simulations with the host model HadGEM; a third variant uses the MT\_CKD 3.2 treatment of the water vapor continuum in lieu of the CAVIAR continuum used in the development of the parameterization.

### 3.2 Instantaneous clear-sky forcing at present day

Figure 2 shows an example calculation of instantaneous radiative forcing: the change in net downward flux at the top-of-atmosphere (TOA) and the surface, and the change in net absorption across the atmosphere (net flux at TOA minus net at surface), here for the change between present-day and pre-industrial conditions. Increased greenhouse gas concentrations in the present day increase the opacity of the atmosphere. In the longwave this acts to decrease outgoing longwave at the TOA and increase downward longwave at the surface. The increase in downwelling surface radiation is smaller than the decrease in outgoing longwave, resulting in decreased radiative cooling across the atmosphere. In the shortwave there a near-zero increase in scattering back to space but an increase in atmospheric absorption, resulting in diminished solar radiation at the surface.

Agreement among the line-by-line models is excellent: the standard deviation across all six quantities (forcing at the top-of-atmosphere, with the atmosphere, and at the surface, for longwave and shortwave) is less than 0.025 W m<sup>-2</sup> with the exception of LW absorption, where the standard deviation is 0.033 W m<sup>-2</sup>. There is no systematic variation across forcing variants, indicating that the equivalent concentrations accurately sum-



**Figure 2.** Global, annual mean instantaneous clear-sky radiative forcing by greenhouse gases at present-day, relative to pre-industrial conditions, as computed by benchmark radiative transfer models. Longwave results are on the left, shortwave results on the right, with the reference model denoted by the color. Model names follow the RFMIP convention with contributions from SOCRATES labeled as HadGEM3 to link the results to the host climate model. Results include multiple representations of greenhouse gas changes (circles, squares, and diamonds corresponding to forcing variants 1, 2, and 3) and small variants in the treatment of some physical processes as explained in the text. All variants of the reference models agree well in longwave calculations, while SOCRATES results in the shortwave show the small but noticeable impact of different treatments of the H<sub>2</sub>O continuum, which overlaps with absorption by other gases in the near-infrared and so affects forcing by those gases.

260 marize the radiative impact of the neglected gases in the transition from pre-industrial  
 261 to present-day conditions.

262 Changes in shortwave flux between pre-industrial and present-day are substantially  
 263 smaller than in the longwave. The standard deviations are commensurate with those in  
 264 the longwave, but diversity in atmospheric absorption and surface forcing is dominated  
 265 by physics variant 2 of the SOCRATES code, which is unique among the models in using  
 266 the CAVIAR treatment for continuum absorption by water vapor (Ptashnik et al.,  
 267 2011, 2013). Absorption in the near infrared in the CAVIAR continuum is substantially  
 268 larger than in the MT\_CKD continuum on which all other models rely, especially where  
 269 water vapor absorption coincides with absorption lines of CO<sub>2</sub>, CH<sub>4</sub> and N<sub>2</sub>O. This masks  
 270 changes in opacity due to well-mixed greenhouse gases and reduces the forcing at the surface  
 271 between pre-industrial and present-day concentrations.

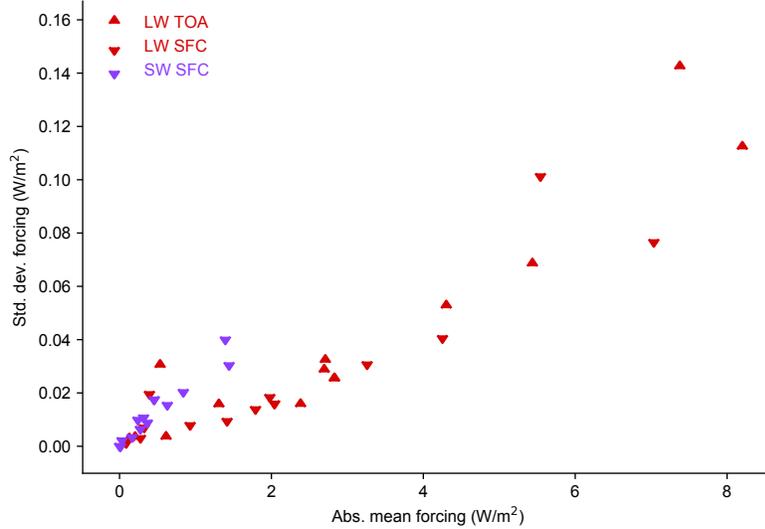
272 Global-mean values of clear-sky instantaneous radiative forcing for a range of well-  
 273 mixed greenhouse gases, averaged across all available reference models, are provided in  
 274 Table 3.2. Variability across models and forcing and physics variants, in both longwave  
 275 and shortwave forcing calculations, increases with the magnitude of the forcing (Figure  
 276 3).

### 277 3.3 Establishing a benchmark for parameterization error

278 Experiment *rad-irf* is intended to assess error in the parameterization of clear-sky  
 279 radiation in the climate models participating in CMIP6. Resolving this error is only possible  
 280 if the disagreement among benchmark models is small relative to the typical difference  
 281 between a parameterization and the reference models themselves. (Sampling er-

**Table 2.** Instantaneous radiative forcing computed as the mean across all available benchmark models, forcing variants, and physics variants, in  $\text{W m}^{-2}$ . Forcing is defined as net downward flux under perturbed conditions minus net downward flux under pre-industrial (PI) conditions; because the profiles provided for experiment *rad-irf* are perturbed around present-day (PD) conditions the difference required may be indirect, as explained in the table. Values are provided for the top of the atmosphere (TOA) and surface (Sfc). RFMIP experiment *rad-irf* contains further perturbations meant to assess errors in temperature and humidity dependence.

Experiment	LW TOA	LW Sfc	SW TOA	SW Sfc
Computed as difference from perturbation “PI”				
Present-day	2.830	2.040	0.055	-0.455
Future	7.377	5.542	0.355	-1.393
Last Glacial Maximum	-2.384	-1.416	-0.065	0.316
Computed as negative difference from perturbation “PD”				
Present-day $\text{CO}_2$	1.308	0.929	0.029	-0.165
Present-day $\text{CH}_4$	0.613	0.275	0.055	-0.242
Present-day $\text{N}_2\text{O}$	0.205	0.088	0.002	-0.011
Present-day $\text{O}_3$	0.129	0.325	-0.032	-0.033
Present-day halocarbons	0.534	0.393	0.000	-0.001
Computed as difference from perturbation “PI $\text{CO}_2$ ”				
$\frac{1}{2}\times\text{CO}_2$	-2.695	-1.790	-0.050	0.274
$2\times\text{CO}_2$	2.709	1.978	0.064	-0.367
$3\times\text{CO}_2$	4.302	3.260	0.110	-0.629
$4\times\text{CO}_2$	5.436	4.252	0.146	-0.840
$8\times\text{CO}_2$	8.201	7.035	0.252	-1.442

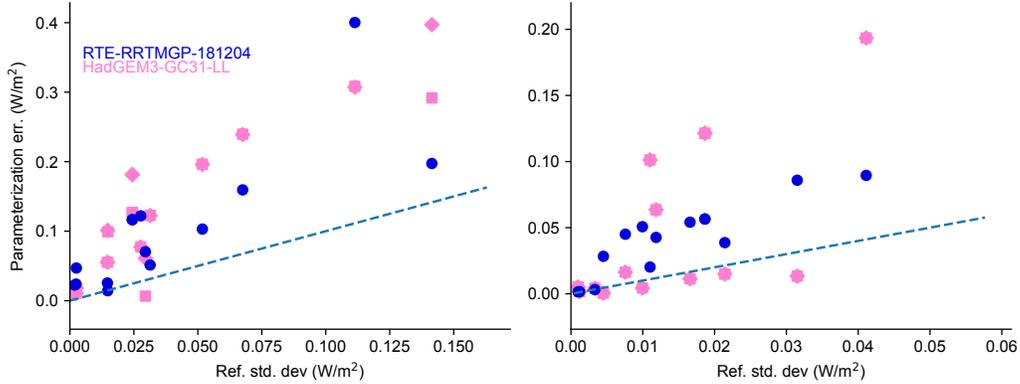


**Figure 3.** Standard deviation in estimates of global-mean instantaneous radiative forcing by greenhouse gases as a function of the absolute value of mean forcing across 18 benchmark calculations in the longwave (red) and nine in the shortwave (purple). Top-of-atmosphere forcing is indicated with an upward-pointing triangle; forcing at the surface with a downward-pointing triangle. Only forcing at the surface is shown for the shortwave. Agreement across models, forcing variants, and model physics variants increases with the mean forcing but it roughly two orders of magnitude smaller than the mean forcing across longwave experiments. Shortwave experiments are a factor of 2-3 more variable, partly driven by different treatments of near-infrared water vapor continuum. The figure illustrates agreement with respect to changed greenhouse gas concentrations; perturbations in experiment *rad-irf* in which temperature and/or humidity changes are omitted.

282 ror is smaller than the difference across reference models; see Figure 1). Figure 4, which  
 283 compares error from two modern parameterizations to the variability across the refer-  
 284 ence models, suggests that the benchmark calculation is likely to meet this goal. Results  
 285 are shown for forcing across all 17 perturbations in experiment *rad-irf*. Errors relative  
 286 to LBLRTM v12.8 are shown the for low spectral-resolution version of SOCRATES, as  
 287 used in the HadGEM model, and for the newly-developed RTE+RRTMGP code (Pin-  
 288 cus et al., 2019) which is trained on calculations with LBLRTM v12.8. These parame-  
 289 terizations use recent spectroscopic information and so are likely to be among the pa-  
 290 rameterizations with the smallest error. Nonetheless the error in each parameterization  
 291 is almost always larger than the standard deviation across reference models, indicating  
 292 differences between parameterizations and all reference models are dominated by param-  
 293 eterization error.

#### 294 4 Towards effective radiative forcing

295 RFMIP experiment *rad-irf* was designed to assess parameterization error but the  
 296 benchmark calculations might also be exploited to refine knowledge of the radiative forc-  
 297 ing experienced by Earth due to various composition changes. Two conceptually differ-  
 298 ent steps are required, both of which are likely to make the estimate substantially less  
 299 certain. One is accounting for the impact of clouds, which requires radiative calculations  
 300 over the large range of imperfectly-characterized cloud properties. The other is account-



**Figure 4.** Absolute error in instantaneous radiative forcing (longwave at the top of atmosphere on the left, shortwave at the surface on the right) as computed by two parameterizations, both based on current spectroscopic information, as a function of amount of disagreement across the reference models. Results are shown for all available forcing and physics variants for each of the 17 perturbations in experiment *rad-irf*. Error is assessed relative to LBLRTM v12.8 on which the RTE+RRTMGP parameterization is trained, minimizing the error for this parameterization. Regardless of which model is used as the benchmark, however, the error in each parameterization exceeds the standard deviation of results from the reference models for a large majority of perturbations, indicating that the reference calculations reported here are accurate enough to resolve parameterization error.

301 ing for adjustments (see Section 1) which introduces conceptually more uncertain non-  
 302 radiative calculations. The long history of efforts to establish high-precision estimate of  
 303 forcing by greenhouse gases (e.g., most recently, Myhre et al., 2006; Etminan et al., 2016)  
 304 provides a point of reference for any efforts to leverage RFMIP calculations.

#### 305 4.1 Accounting for clouds

306 Clouds modulate radiative forcing by greenhouse gases by screening changes in con-  
 307 centration behind the cloud. The degree to which clouds obscure greenhouse gas forc-  
 308 ing depends primarily on the cloud optical depth (though longwave emissivity and short-  
 309 wave reflectance and transmittance). Top of atmosphere forcing is also modulated by  
 310 surface properties and cloud top height or pressure; surface forcing is modulated by cloud  
 311 base height. Accounting for clouds in estimates of radiative forcing by greenhouse gases  
 312 requires characterizing the wide variation in these properties in space and time. Obser-  
 313 vations from passive satellite sensors offer the best sampling of global variations but pro-  
 314 vide much stronger constraints on the quantities that affect top-of-atmosphere forcing  
 315 than surface forcing. Previous efforts to establish benchmarks for radiative forcing (e.g.  
 316 Etminan et al., 2016; Myhre et al., 2006) have used two atmospheric profiles (see Sec.  
 317 2) each combined with three sets of representative cloud properties as observed by pas-  
 318 sive satellite instruments. Sampling errors in the global, annual mean at the top of the  
 319 atmosphere are thought to be of order 1% although this error estimate has not been re-  
 320 visited since the 1990s (Myhre & Stordal, 1997; Freckleton et al., 1998). Errors in cloud  
 321 impacts on surface forcing have not been assessed.

322 We hope to revisit this question in future work. One important question will be  
 323 whether computational effort is better spent in sampling the co-variability of cloud prop-  
 324 erties with other atmospheric and surface properties or in high-spectral resolution cal-

**Table 3.** Ratio of all-sky to clear-sky instantaneous radiative forcing, at the top-of-atmosphere and the surface, across a range of models and experiments in CMIP6. Clear-sky and all-sky (including clouds) fluxes are computed using a second radiative transfer calculation in which the forcing agents are modified for diagnostics purposes. Results from HadGEM3 and IPSL-CM6A use diagnostic calculations requested for CFMIP in which CO<sub>2</sub> concentrations are quadrupled from pre-industrial values. Values from GFDL-CM4, performed for this work, are computed by setting forcing agents to pre-industrial values in three RFMIP fixed-SST integrations. Results from HadGEM3 are preliminary and may be revised before they are made publicly available. Shortwave forcing at the top of atmosphere is so small that inferences of cloud masking are quite uneven across models.

experiment	HadGEM3-GC31-LL	IPSL-CM6A-LR	GFDL-CM4		
	amip	historical	4xCO2	GHG	anthro
LW TOA	0.764	0.735	0.763	0.757	0.767
LW SFC	0.622	0.608	0.696	0.689	0.680
SW SFC	0.718	0.732	0.711	0.853	0.714

325 calculations to limit approximation errors. These questions, though, are beyond the scope  
 326 of what can be accomplished with reference model calculations to *rad-irf*. As an alter-  
 327 native we have examined the ratio of all-sky to clear-sky instantaneous radiative forc-  
 328 ing by greenhouse gases in the few available simulations from CMIP6. The Cloud Feed-  
 329 backs Model Intercomparison Project (Webb et al., 2017) requests, at low priority, cal-  
 330 culations with CO<sub>2</sub> concentrations quadrupled from pre-industrial concentrations; two  
 331 models have made such calculations available at this writing (HadGEM3 for experiment  
 332 *amip* and IPSL-CM6A for experiment *historical*). We have also made diagnostic radi-  
 333 ation calculations in GFDL’s AM4 model (Zhao et al., 2018) using pre-industrial green-  
 334 house gas concentrations during RFMIP “fixed-SST” experiments in which these con-  
 335 centrations are normally held constant at present-day values; these follow the protocol  
 336 described by (Lin et al., 2017).

337 Results are provided in Table 4.1. A decade ago Andrews & Forster (2008) found  
 338 that the presence of clouds reduced longwave instantaneous radiative forcing from quadru-  
 339 pled CO<sub>2</sub> concentrations by amounts ranging from 9 to 20%, depending on the model  
 340 (see their Table S2). As the distribution of clouds simulated by climate models has con-  
 341 tinued to move closer to observations (e.g. Klein et al., 2013) the estimated impact on  
 342 top-of-atmosphere forcing has grown while the range across models and experiments has  
 343 decreased (in Table 4.1 it is 23.6% to 26.5%). Clouds have a similar impact on short-  
 344 wave forcing at the surface and an even larger impact on longwave forcing at the sur-  
 345 face, though weaker observational constraints on the vertical structure of clouds allow  
 346 for greater diversity across models.

## 347 4.2 Accounting for adjustments from temperature changes in the strato- 348 sphere

349 As explained in Section 1 the measure of forcing most closely related to temper-  
 350 ature response is effective radiative forcing: the sum of the instantaneous radiative forc-  
 351 ing, computable with robust radiative transfer models, and adjustments made by the phys-  
 352 ical climate system in the absence of surface temperature change (Sherwood et al., 2015).  
 353 Adjustments, like forcing, result from a difference in two states and so are not directly  
 354 observable. Many adjustments involve changes to circulations and clouds across a range

of scales (e.g. Gregory & Webb, 2008; Bretherton et al., 2013; Merlis, 2015) and can only be assessed with dynamical models for which establishing benchmarks is impractical.

In the climate models used to assess the global magnitude and distributions of adjustments, the dominant adjustment to greenhouse gas forcing is consistently the cooling of the stratosphere, partly because various tropospheric adjustments counteract each other (e.g. Smith et al., 2018; Smith, Kramer, & Sima, 2020). This cooling was first noted by Manabe & Wetherald (1967) and identified as an adjustment to longwave forcing by Hansen et al. (1997). As Shine & Myhre (2020) explain, increased concentrations of well-mixed greenhouse gases increase both emission by the stratosphere and absorption of radiation emitted from the troposphere. If the background atmosphere is optically thick in the spectral region in which the gas is active (e.g. for CO<sub>2</sub>) additional warming from tropospheric emission is small and the stratosphere cools, enhancing instantaneous forcing at the top of the atmosphere, but if the the background atmosphere is optically thin (as for most halocarbons) the stratosphere may warm, damping the instantaneous forcing.

The magnitude of this adjustment can be computed to a good approximation by assuming that dynamical heating in the stratosphere is fixed (Ramanathan & Dickinson, 1979; Fels et al., 1980): computing the radiative cooling rate of the stratosphere under baseline (present-day) conditions, assuming that this cooling is balanced by dynamical heating, and then finding the temperature profile necessary to obtain the same net cooling profile under changed greenhouse gas concentrations. We follow Myhre et al. (2006) and Etminan et al. (2016) in supplying this first-order estimate of adjustments. We compute the adjustment caused by stratospheric temperature re-equilibration, assuming fixed dynamical heating, by iterating with GRTCODE model at reduced spectral resolution until radiative heating rates reach their values in the present-day atmosphere. The calculations assume a uniform tropopause pressure of 200 Pa and account for changes in both longwave and shortwave heating rates. For well-mixed greenhouse gases the impact of stratospheric temperature adjustment depends primarily on the spectral region in which the gas absorbs.

The impact of stratospheric temperature adjustment, expressed as the ratio of the change in flux due to temperature equilibration to the instantaneous longwave radiative forcing, is shown for a range of concentrations at present-day relative to pre-industrial conditions in Table 4.2. Stratospheric temperature changes from well-mixed greenhouse gases amplify (CO<sub>2</sub>, N<sub>2</sub>O) or damp (CH<sub>4</sub>, halocarbons) forcing at the top of the atmosphere; for all gases but CO<sub>2</sub> the impact is just a few percent. Surface forcing is damped by a similar amount.

Carbon dioxide is a notable exception: the amplification of top-of-atmosphere forcing at present-day is more than 55%. This large adjustment occurs because the total forcing at the top-of-the-atmosphere is a balance between contributions from distinct spectral regions. Near the center of the 15  $\mu\text{m}$  absorption band of CO<sub>2</sub> the atmosphere is optically thick and emission to space occurs in the stratosphere; increases CO<sub>2</sub> concentrations tends to increase outgoing longwave radiation because stratospheric temperature increases with height. Away from the band center the atmosphere is optically thin, emission is from the troposphere, and increasing concentrations acts to decrease outgoing longwave radiation. Net forcing is negative (see Table 3.2) because the the tropospheric contribution dominates. Stratospheric cooling damps the instantaneous forcing from the band center, allowing the optically-thin regions to dominate the change in top-of-atmosphere flux even more effectively. The adjustment also increases by 1.8% per  $\text{W m}^{-2}$  (Figure 5) so that effective radiative forcing may be modestly super-logarithmic in CO<sub>2</sub> concentrations even though the instantaneous radiative forcing is nearly perfectly logarithmic.

Stratospheric temperature adjustment nearly doubles the top-of-atmosphere instantaneous forcing from ozone but for quite different reasons. Ozone concentrations at present-

**Table 4.** Ratio of adjustment due to stratospheric temperature equilibration under the fixed dynamical heating assumption to instantaneous clear-sky longwave radiative forcing at the top of atmosphere and the surface for a range of forcing agents. Both forcing and stratospheric adjustment are computed using GFDL GRTCODE line-by-line model. Shortwave adjustments are all essentially zero.

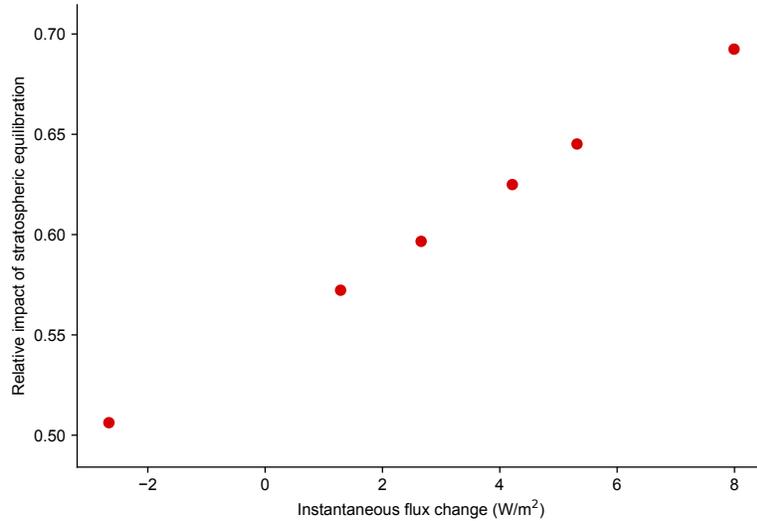
Experiment	TOA	SFC
Present-day	0.31	-0.03
Present-day CO <sub>2</sub>	0.57	-0.05
Present-day CH <sub>4</sub>	-0.05	0.01
Present-day N <sub>2</sub> O	0.03	-0.01
Present-day O <sub>3</sub>	1.90	-0.06
Present-day halocarbons	-0.11	0.01

407 day vary substantially in the vertical, peaking in the stratosphere. As one consequence  
 408 ozone acts to heat the stratosphere near the center of the 10  $\mu\text{m}$  band and increases in  
 409 ozone concentration in either the troposphere or stratosphere tend to decrease net radi-  
 410 ation at the top of the atmosphere. The vertical distribution of change is also non-uniform:  
 411 relative to pre-industrial conditions ozone concentrations have increased in the tropo-  
 412 sphere but decreased in the stratosphere. The modest positive forcing from present-day  
 413 ozone relative to pre-industrial conditions results from a slightly larger decrease in out-  
 414 going radiation from tropospheric emission than can be balanced by increased emission  
 415 from concentration reductions in the stratosphere. The stratosphere cools modestly de-  
 416 spite because reduced concentrations of ozone because decreases in absorption of short-  
 417 wave radiation are larger than the increases from enhanced longwave emission. This cool-  
 418 ing, too, reduces the stratospheric contribution to forcing. Stratospheric adjustment of  
 419 ozone is larger than for carbon dioxide, in a relative sense, only because the balance be-  
 420 tween stratosphere and troposphere is more even for instantaneous forcing.

## 421 5 Constraints on radiative forcing

422 Previous work (e.g. Chung & Soden, 2015; Soden et al., 2018) has established that  
 423 the instantaneous radiative forcing for a given change in atmospheric composition can  
 424 vary widely among climate models. This diversity has two distinct sources: parameter-  
 425 ization error and variety in the distributions of temperature, humidity, and clouds be-  
 426 tween models. By using accurate models across a representative set of observed condi-  
 427 tions we have shown that the true value of clear-sky instantaneous radiative forcing can  
 428 be determined quite precisely, with all-sky estimates limited primarily by challenges in  
 429 representing the co-variability of clouds and atmospheric state. This highlights the dis-  
 430 tinction between climate model diversity and true uncertainty in estimates of instantane-  
 431 ous radiative radiative forcing. Adjustments arising from greenhouse gas forcing, how-  
 432 ever, remain a currently-irreducible source of uncertainty in attempts to estimate the true  
 433 effective radiative forcing to which our planet has been subject and a source of poorly-  
 434 constrained diversity among model estimates of effective radiative forcing.

435 Two caveats apply to our estimates of clear-sky instantaneous radiative forcing. First,  
 436 RFMIP explores parameterization error in perturbations around present-day conditions,  
 437 so that our estimates of instantaneous radiative forcing are based on present-day distri-  
 438 butions of temperature and humidity. Forcing depends modestly on both quantities (Huang  
 439 et al., 2016) so our estimates of forcing are slightly enhanced relative to calculations that  
 440 use pre-industrial conditions. Second, in the interests of highlighting model error in the  
 441 representation of absorption by gases, the *rad-irf* protocol specifies spectrally-constant



**Figure 5.** Ratio of stratospheric temperature adjustment to instantaneous radiative forcing at the top of the atmosphere for CO<sub>2</sub> perturbations ranging from 0.5× to 8× pre-industrial concentrations. Assuming that heating from atmospheric dynamic stays constant allows the computation of a new equilibrium temperature profile to be computed; this profile is colder (because the stratosphere is a more effective emitter) so the adjustment amplifies instantaneous radiative forcing. The magnitude of the adjustment depends modestly on the magnitude of the forcing itself, suggesting that effective radiative forcing by CO<sub>2</sub> may be slightly super-logarithmic in concentration even if the instantaneous radiative forcing is not.

442 surface albedo and emissivity as obtained from ERA-Interim. Shortwave forcing at the  
 443 top of the atmosphere, which arises from the sensitivity to greenhouse gases of radiation  
 444 reflected at the surface and transmitted through the atmosphere, can be dramatically  
 445 overestimated if the surface albedo is overestimated in the spectral range affected by a  
 446 given gas (Oreopoulos et al., 2012). The small values of shortwave forcing in Table 3.2  
 447 suggest that the simple treatment of surface albedo is not likely to cause a large error  
 448 but accounting for spectral variations in surface albedo would be a useful exercise.

449 The agreement in global-mean instantaneous radiative among reference models, though  
 450 encouraging, is consistent with almost 30 years of experience: Ellingson et al. (1991), for  
 451 example, report that most of their line-by-line results for flux agree to within 1%. The  
 452 agreement arises partly because radiative forcing, as the difference between two calcu-  
 453 lations, is also less sensitive to assumptions or subtle differences between models because  
 454 many variations cancel out (Mlynczak et al., 2016). In our data set, however, the level  
 455 of agreement in fluxes across models at the atmosphere’s boundaries under present-day  
 456 conditions varies by less than 0.6 W m<sup>-2</sup> in the longwave and 0.7 W m<sup>-2</sup> in the short-  
 457 wave - smaller than the variability in forcing estimates, in a relative sense, by an order  
 458 of magnitude. The agreement in both fluxes and forcing arises because the models rely  
 459 on the same underlying physics applied to small variants around the same spectroscopic  
 460 data, so that the accuracy is limited by current spectroscopic knowledge more than by  
 461 the ability to calculate fluxes from that knowledge. So while spectroscopic knowledge  
 462 is now demonstrably more complete than it was 30 years ago (Mlawer & Turner, 2016),  
 463 small variations in forcing estimates – high precision – should be understood as being  
 464 conditioned on this knowledge rather than evidence of true accuracy.

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## Data availability

All results for RFMIP experiment *rad-irf* are available on the Earth System Grid Federation (searching for the experiment name is an effective way to find the data). Python scripts and Jupyter notebooks to produce the paper are available at <https://github.com/RobertPincus/rfmip-benchmark-paper-figures> and will be archived at Zenodo, with a DOI, on acceptance. ERA-Interim data were obtained from <https://www.ecmwf.int/en/forecasts/datasets/archive-datasets/reanalysis-datasets/era-interim>. SOCRATES is available from <https://code.metoffice.gov.uk/trac/socrates> under an open source license but requires a free account from the UK Met Office to access the website. Preliminary data for Table 4.1 were provided by Tim Andrews and Alejandro Bodas-Salcedo of the UK Met Office but will be derivable through data provided on the Earth System Grid.

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